

Motivations and Discovery Prospects for Elusive Higgs Boson(s)

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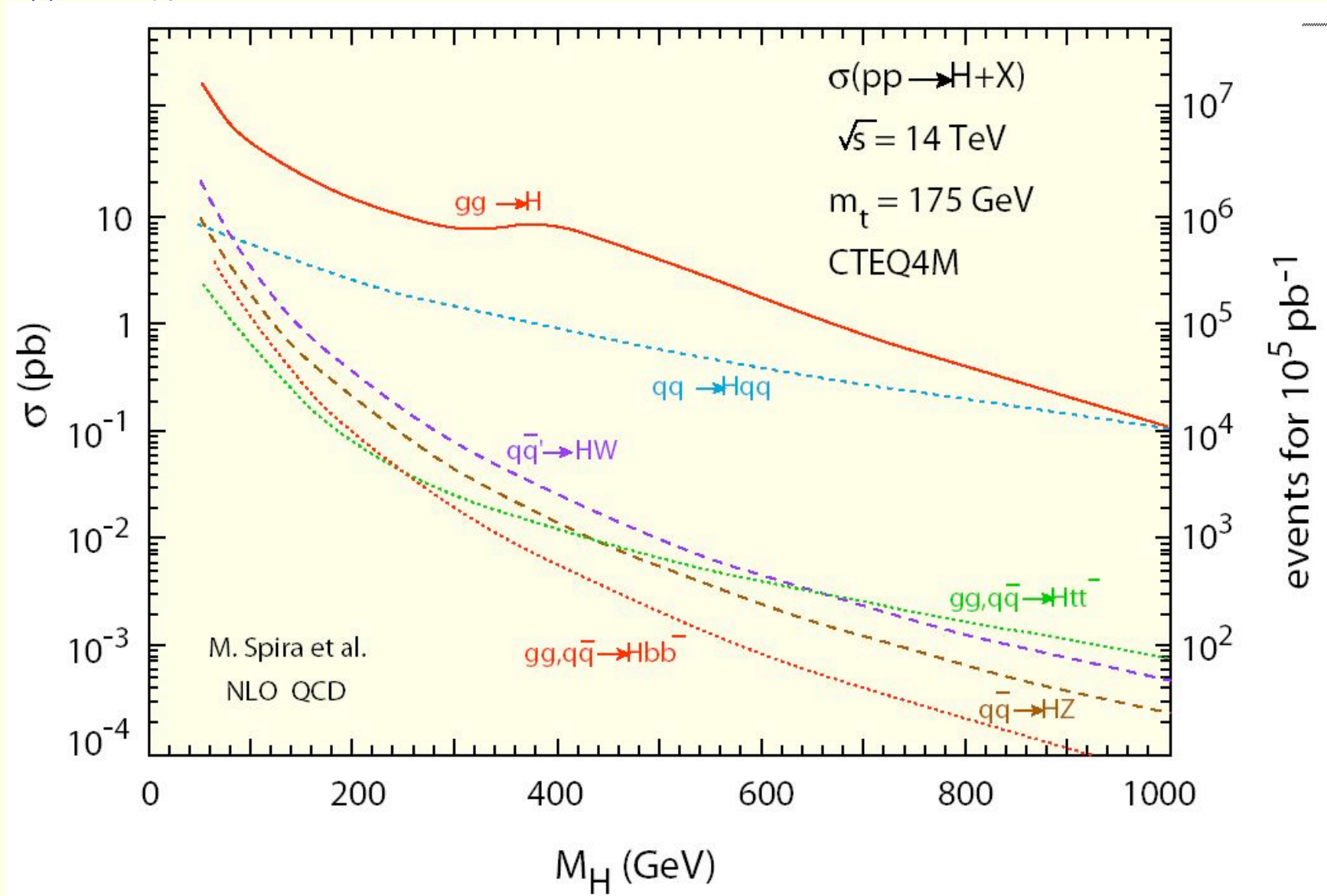
LAPTh, Annecy, Nov. 13, 2009

Synopsis/Outline

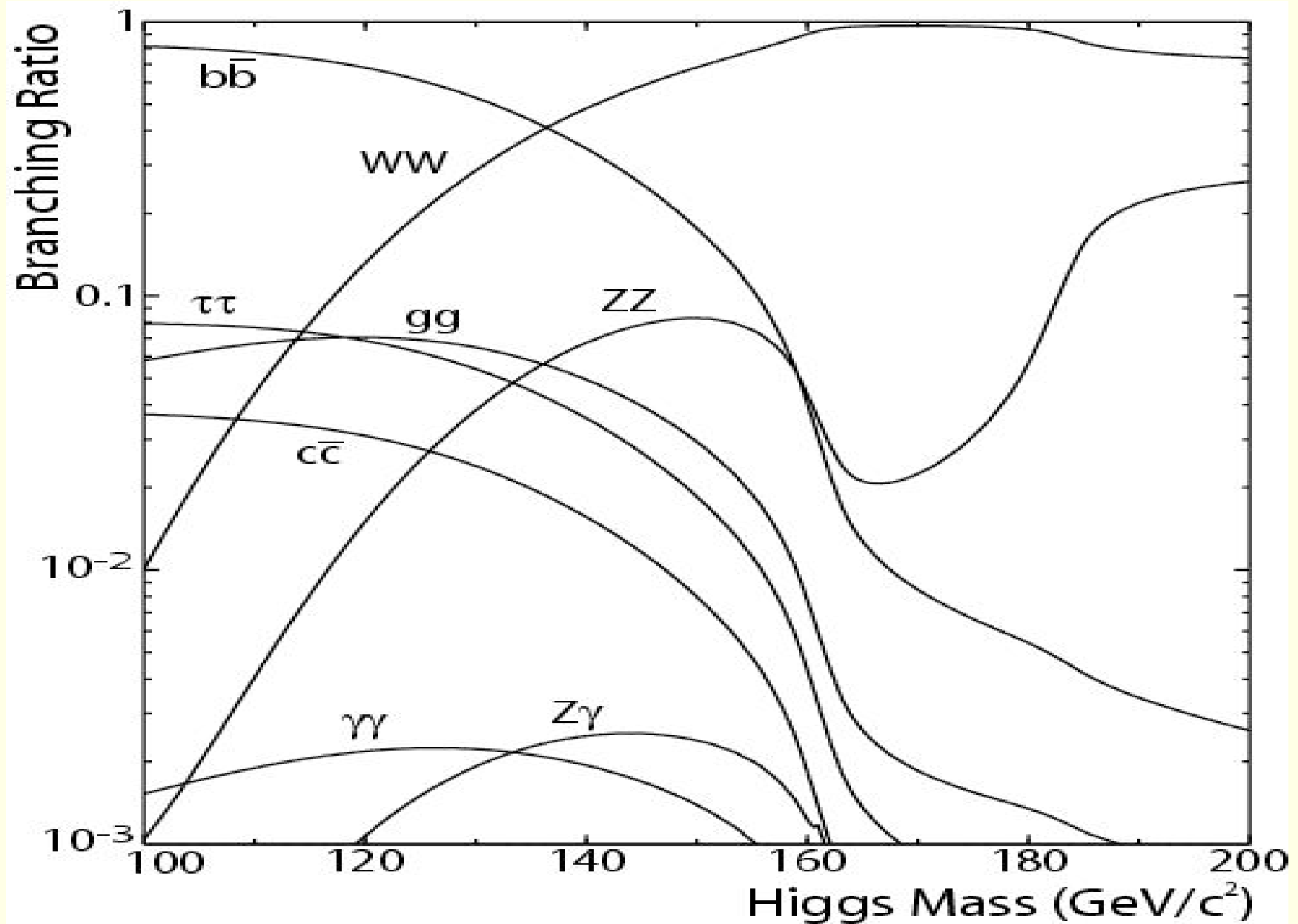
There are excellent motivations for an $m_H \lesssim 105$ GeV **SUSY** Higgs with SM-like couplings to SM particles but elusive decays.

- Precision Electroweak (PEW) data prefer a Higgs boson with SM-like g_{WW_h, ZZ_h} and $m_h \lesssim 105$ GeV
- The simplest solution to the hierarchy problem is SUSY.
- Gauge coupling unification prefers something close to the MSSM.
- Absence of EWSB fine-tuning requires a light SUSY spectrum (in particular, a light \tilde{t}) and a light \tilde{t} implies that the SM-like Higgs of SUSY is light.
- MSSM scenarios having a Higgs with SM-like properties that is light, i.e. $m_h \lesssim 105$ GeV (for PEW perfection) are excluded by LEP.
- Extended SUSY models, including the NMSSM (which preserves all good MSSM features and solves the μ problem) give elusive decay scenarios not ruled out by LEP for $m_h < 105$ GeV.
- LHC strategies for Higgs searches will need to be expanded.

- Higgs cross sections (initiated by SM particles with SM-like h couplings) are determined. Main processes are $gg \rightarrow h$ and $qq \rightarrow q'q'WW$ with $WW \rightarrow h$.



- In the absence of new physics, Higgs decays are also determined by these same couplings.



Or, you could write a letter to the Higgs boson:

Dear Higgs Boson,

We know you're out there. We can feel you now. We know that you're afraid. You're afraid of us; you're afraid of change. We don't know the future. We didn't write this to tell you how this is going to end. We wrote this to tell you how it's going to begin.

As you know, our Large Hadron Collider has had some setbacks due to a.... uh... "transformer malfunction" but we know it was you. You sabotaged our machine. We hope you've been enjoying your vacation because we're scheduled to restart in September 2009 and we're pissed.

....so run and hide, asshole. Run and hide. If you should get careless and allow yourself to get detected by the Tevatron, we are going to be supremely disappointed; because we want to find you first, and when we do, rest assured we are not going to publish right away. We're going to teach you some manners first.

Love,

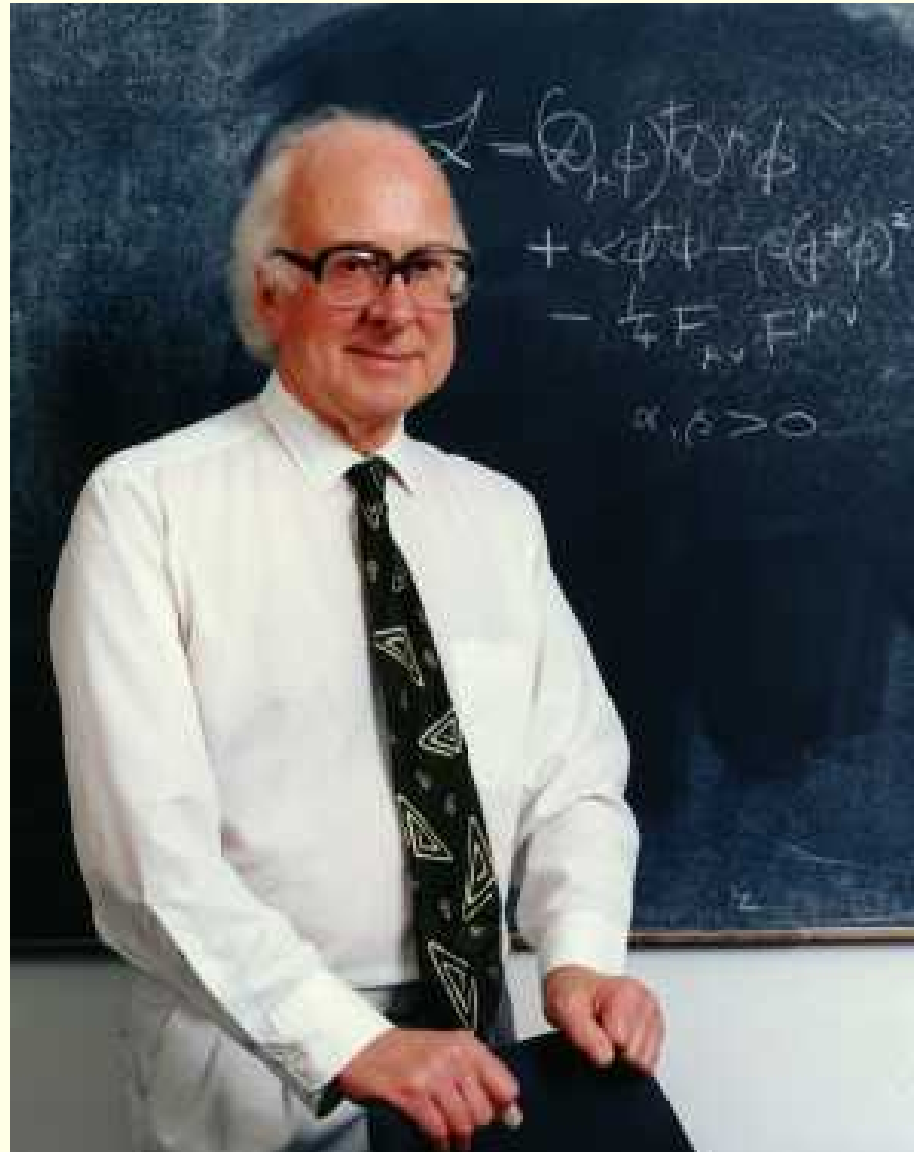
CERN



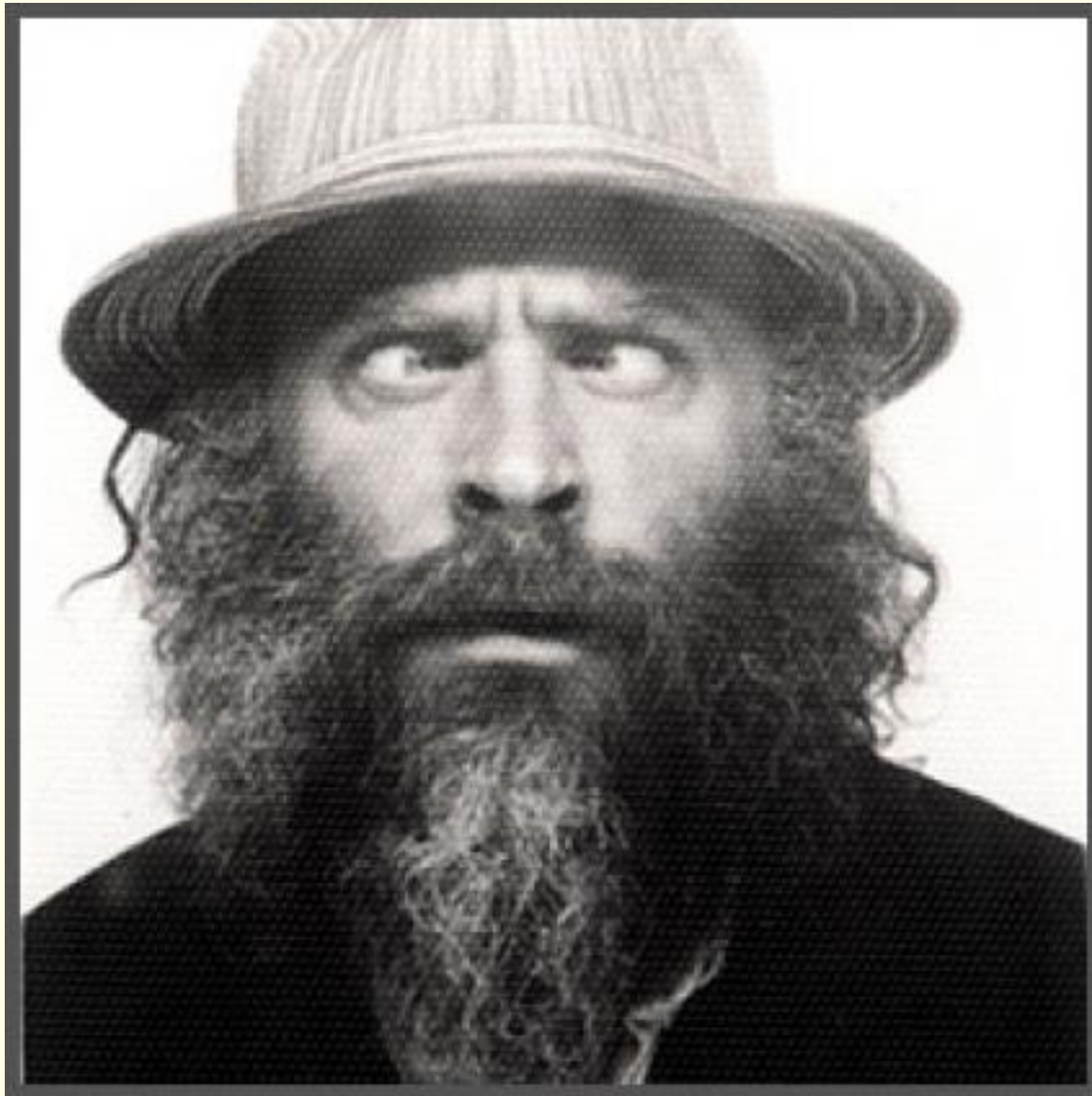
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We really should not count on knowing what the Higgs “looks like”. It could be ...

Priestly, highly orthodox



Ornery/ mean, highly heretical



singer Daniel Higgs

Beautiful but unorthodox

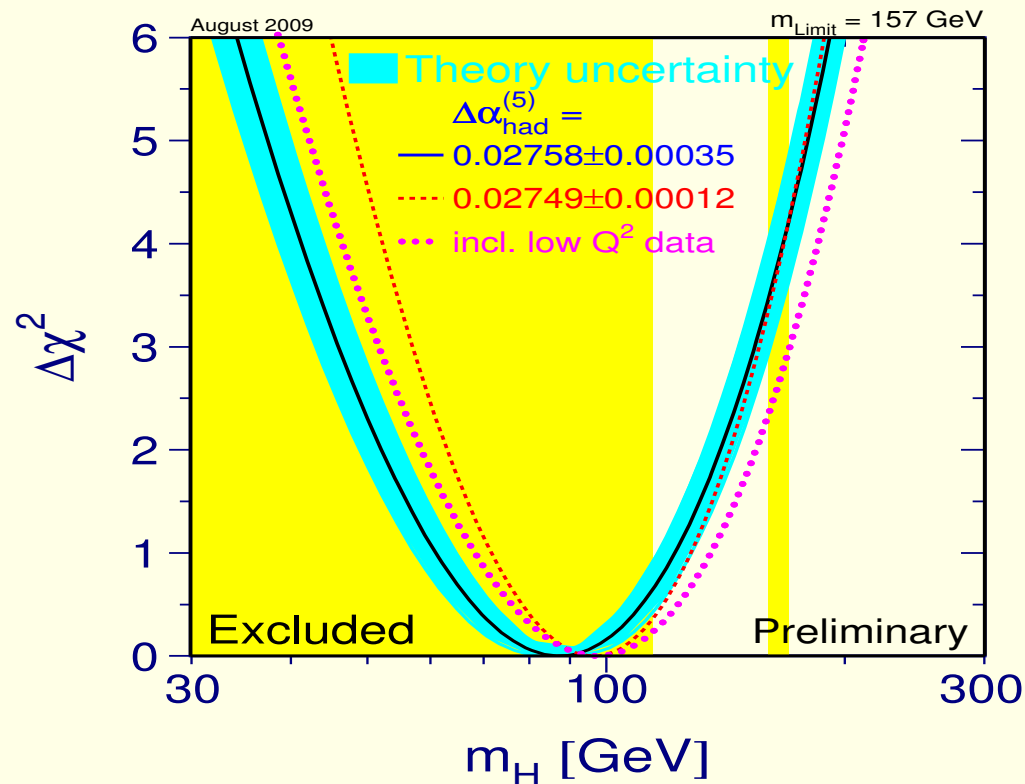


Or, will the LHC bury the Higgs?



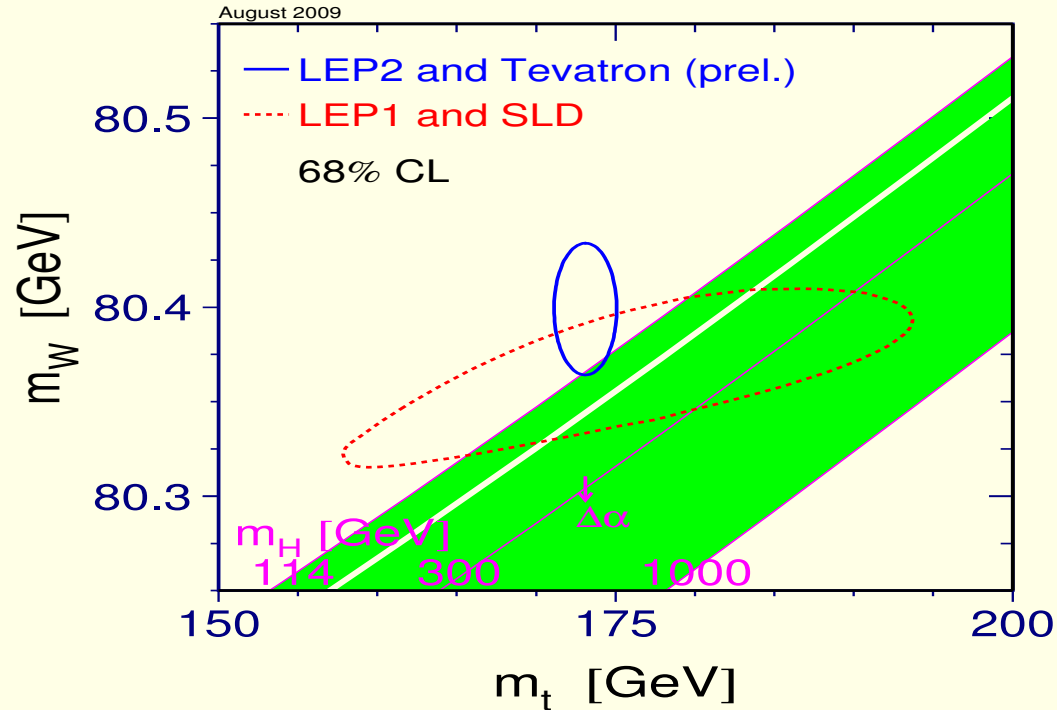
Motivation for Non-Standard Decays — single H

- A fairly recent plot of $\Delta\chi^2(PEW)$ vs. m_H is:



At 95% CL, $m_{h_{\text{SM}}} < 157 \text{ GeV}$ and the $\Delta\chi^2$ minimum is near 85 GeV when all data are included.

The latest m_W and m_t measurements also prefer $m_{h_{SM}} \lesssim 100$ GeV.



Further, the blue-band plot may be misleading due to the discrepancy between the "leptonic" and "hadronic" measurements of $\sin^2 \theta_W^{eff}$, which yield $\sin^2 \theta_W^{eff} = 0.23113(21)$ and $\sin^2 \theta_W^{eff} = 0.23222(27)$, respectively. The SM has a CL of only 0.14 when all data are included.

If only the leptonic $\sin^2 \theta_W^{eff}$ measurements are included, the SM gives a fit with CL near 0.78. However, the central value of $m_{h_{SM}}$ is then near 50 GeV with a 95% CL upper limit of ~ 105 GeV (Chanowitz, xarXiv:0806.0890).

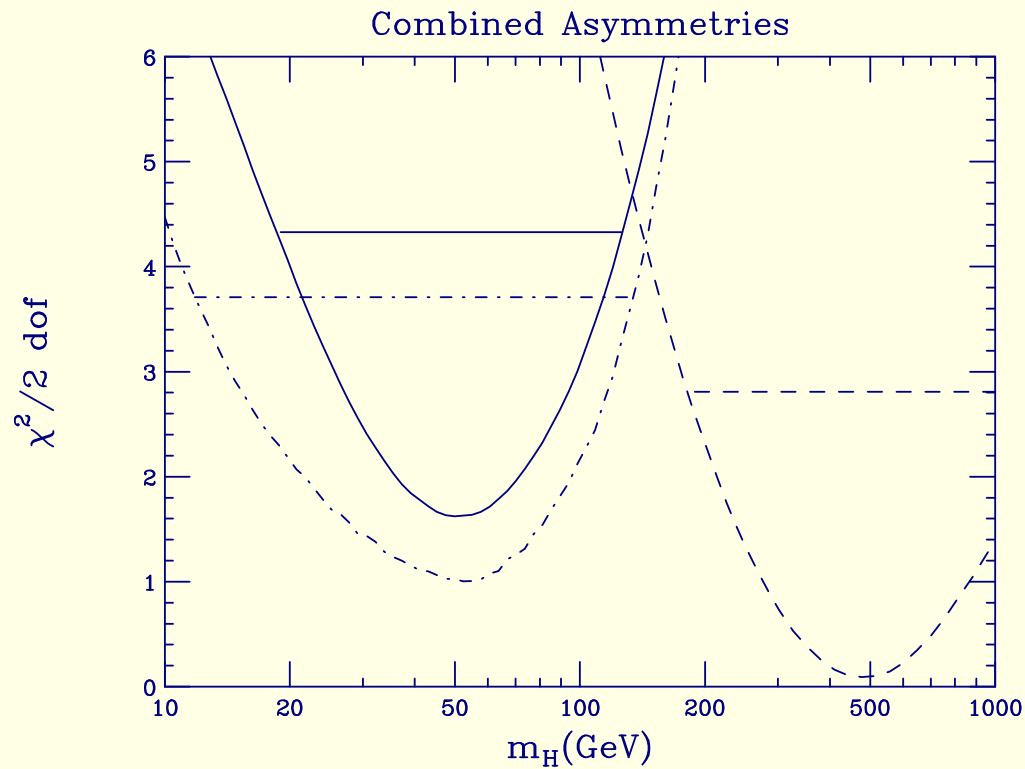


Figure 1: χ^2 distributions as a function of m_H from the combination of the three leptonic asymmetries A_{LR} , A_{FB}^ℓ , $A_\ell(P_\tau)$ (solid line); the three hadronic asymmetries A_{FB}^b , A_{FB}^c , and Q_{FB} (dashed line); and the three m_H -sensitive, nonasymmetry measurements, m_W , Γ_Z , and R_l (dot-dashed line). The horizontal lines indicate the respective 90% symmetric confidence intervals.

- Thus, in an ideal model, a Higgs with SM-like ZZ coupling should have mass no larger than 105 GeV.

But, at the same time, the H must escape LEP and CDF/D0 limits on m_H . In the case of a completely SM-like Higgs they are summarized as

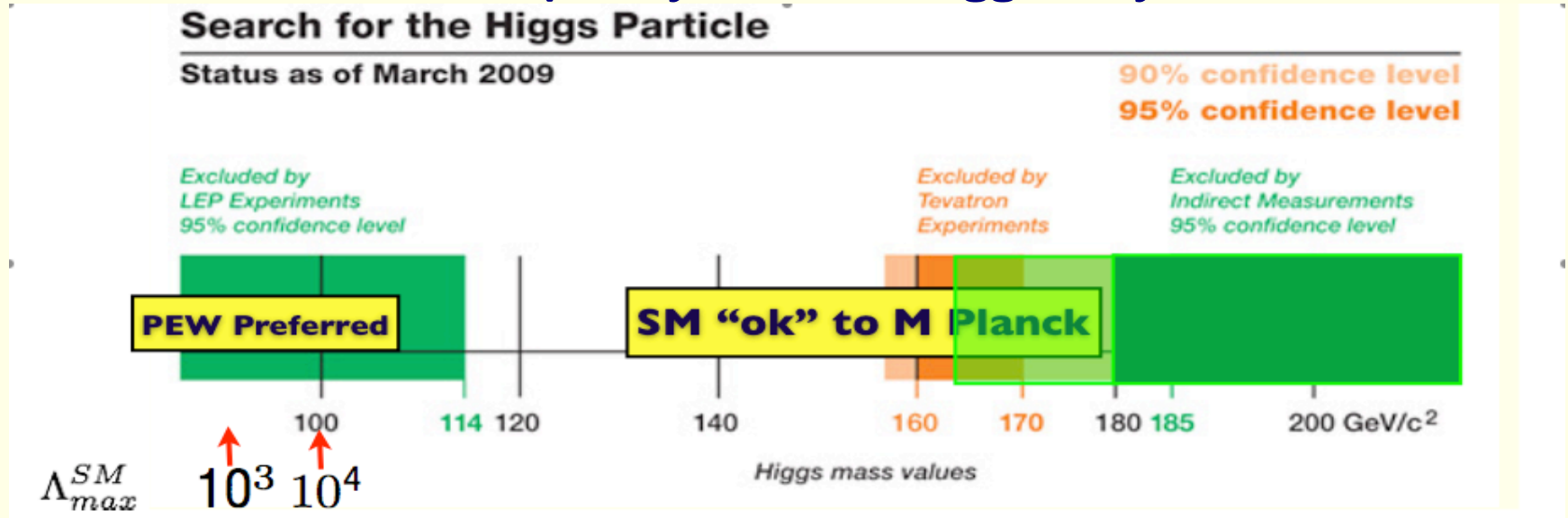


Table 1: LEP m_H Limits for a H with SM-like ZZ coupling, but varying decays. See (S. Chang, R. Dermisek, J. F. Gunion and N. Weiner, Ann. Rev. Nucl. Part. Sci. 58, 75 (2008) [arXiv:0801.4554 [hep-ph]]).

Mode Limit (GeV)	SM modes 114.4	2τ or $2b$ only 115	$2j$ 113	$WW^* + ZZ^*$ 100.7	$\gamma\gamma$ 117	\cancel{E} 114	$4e, 4\mu, 4\gamma$ 114?
Mode Limit (GeV)	$4b$ 110	pure 4τ $86 \rightarrow \sim 105^1$	any (e.g. $4j$) 82	$2f + \cancel{E}$ 90?			

1. Latest ALEPH result.

To have $m_H \leq 105$ GeV requires one of the final three modes.

- Perhaps the ideal Higgs should be such as to predict the 2.3σ excess at $M_{b\bar{b}} \sim 98$ GeV seen in the $Z + b\bar{b}$ final state.

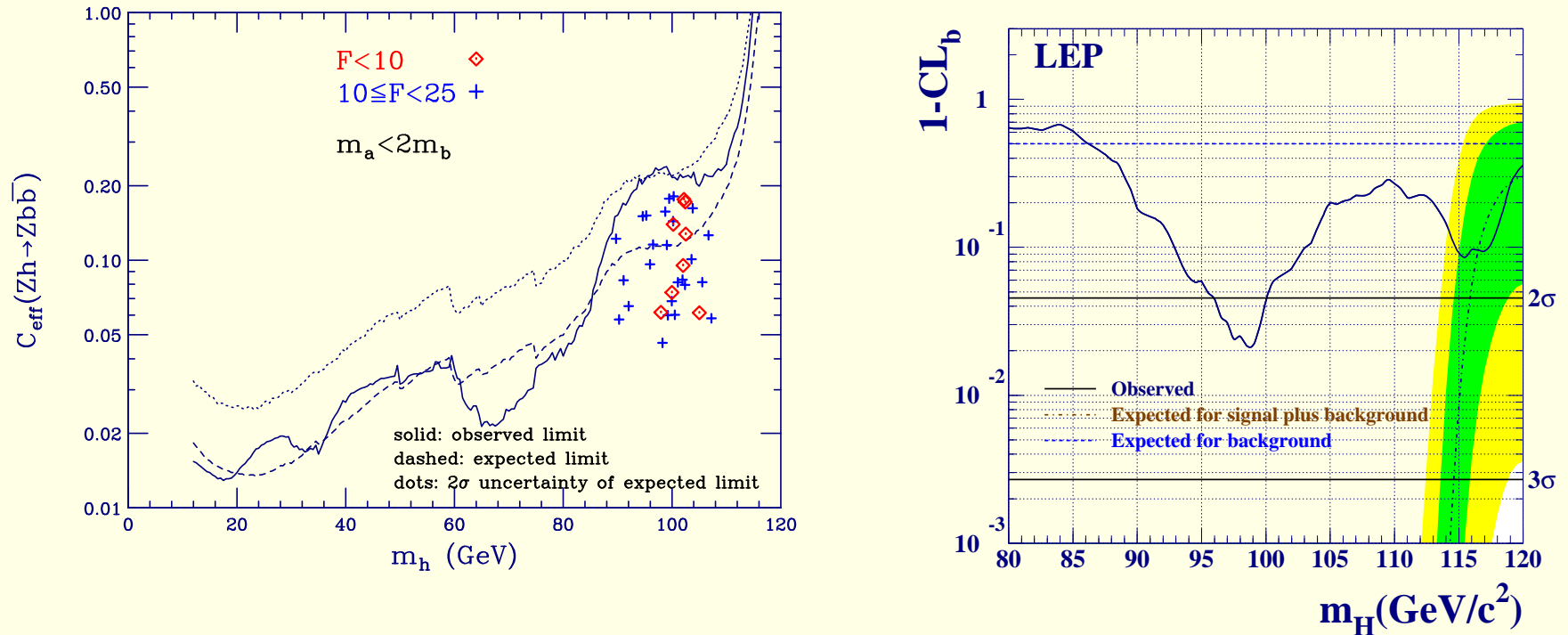


Figure 2: Plots for the $Zb\bar{b}$ final state. F is the m_Z -fine-tuning measure for the NMSSM.

The simplest possibility for the excess is to have $m_H \sim 100$ GeV and $B(H \rightarrow b\bar{b}) \sim (0.1 - 0.2) \times B(H \rightarrow b\bar{b})_{SM}$ (assuming H has SM ZZ coupling as desired for precision electroweak) with the remaining H decays being to one or more of the poorly constrained channels.

- One generic way of having a low LEP limit on m_H is to suppress the $H \rightarrow b\bar{b}$ branching ratio by having a light a (or h) with $B(H \rightarrow aa) > 0.7$ and $m_a < 2m_b$ (to avoid LEP $Z + 4b$ limit at 110 GeV, i.e. above ideal).
For $2m_\tau < m_a < 2m_b$, $a \rightarrow \tau^+\tau^-$. For $m_a < 2m_\tau$, $a \rightarrow jj$.

See: (R. Dermisek and J. F. Gunion, Phys. Rev. Lett. 95, 041801 (2005) [arXiv:hep-ph/0502105]; R. Dermisek and J. F. Gunion, Phys. Rev. D 73, 111701 (2006) [arXiv:hep-ph/0510322])

- **Since the $Hb\bar{b}$ coupling is so small, very modest Haa coupling suffices.**

Higgs pair modes can easily dominate until we pass above the WW threshold.

- So, let us suppose that we want $m_H < 105$ GeV. We should then recall the triviality and global minimum constraints on the scale Λ of new physics.

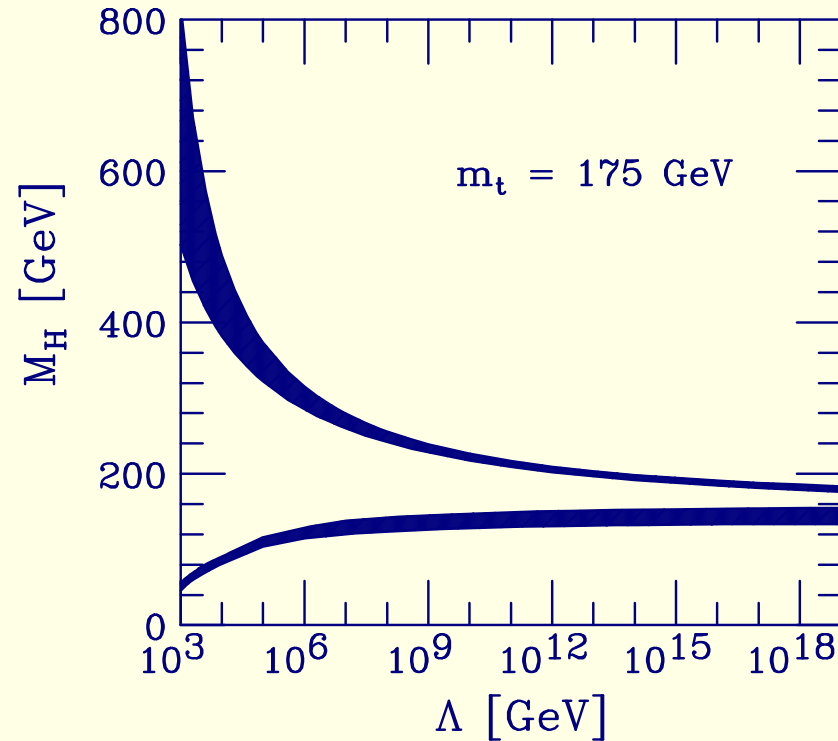


Figure 3: Triviality and global minimum constraints on $m_{h_{\text{SM}}}$ vs. Λ .

The implication is that some new physics should arise for $\Lambda < 10^4 (10^3)$ GeV if $m_h \sim 100$ GeV (~ 50 GeV). A wonderful choice would be SUSY.

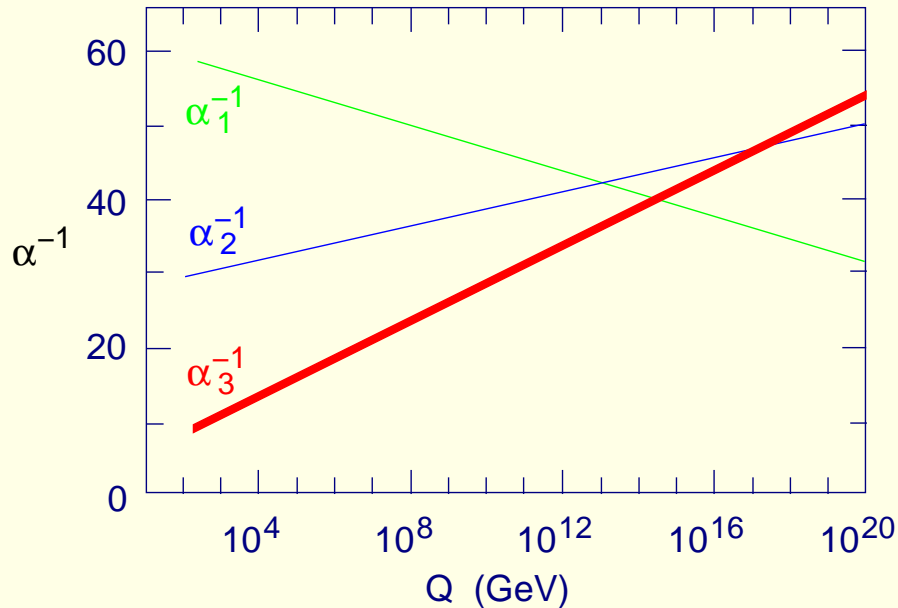
- SUSY does many wonderful things. In particular, SUSY cures the naturalness / hierarchy problem.

- Indeed, the MSSM comes close to being very nice.

If we assume that all sparticles reside at the $\mathcal{O}(1 \text{ TeV})$ scale and that μ is also $\mathcal{O}(1 \text{ TeV})$, then, the MSSM has two particularly wonderful properties.

1. **Gauge Coupling Unification**

Standard Model



MSSM

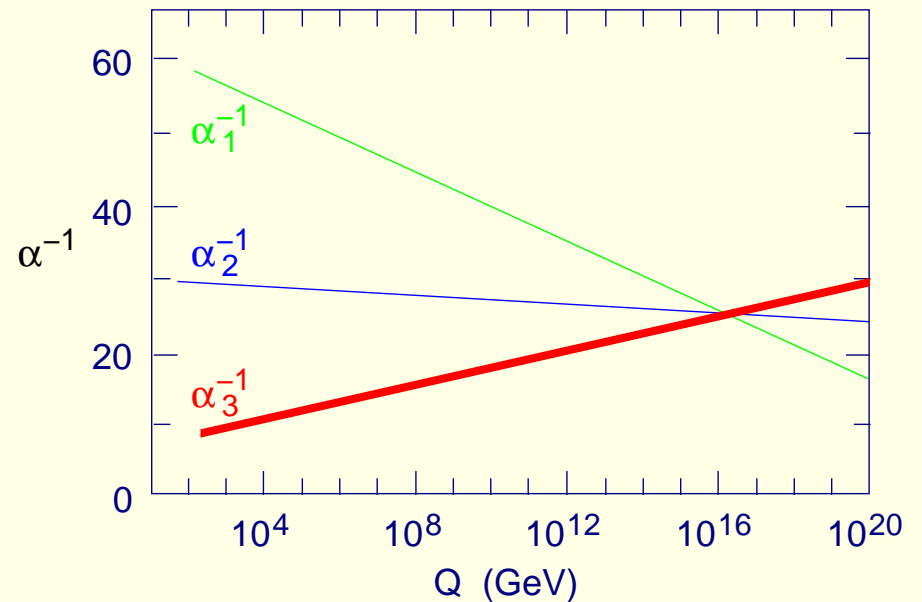


Figure 4: Unification of couplings constants ($\alpha_i = g_i^2/(4\pi)$) in the minimal supersymmetric model (MSSM) as compared to failure without supersymmetry.

2.

RGE EWSB

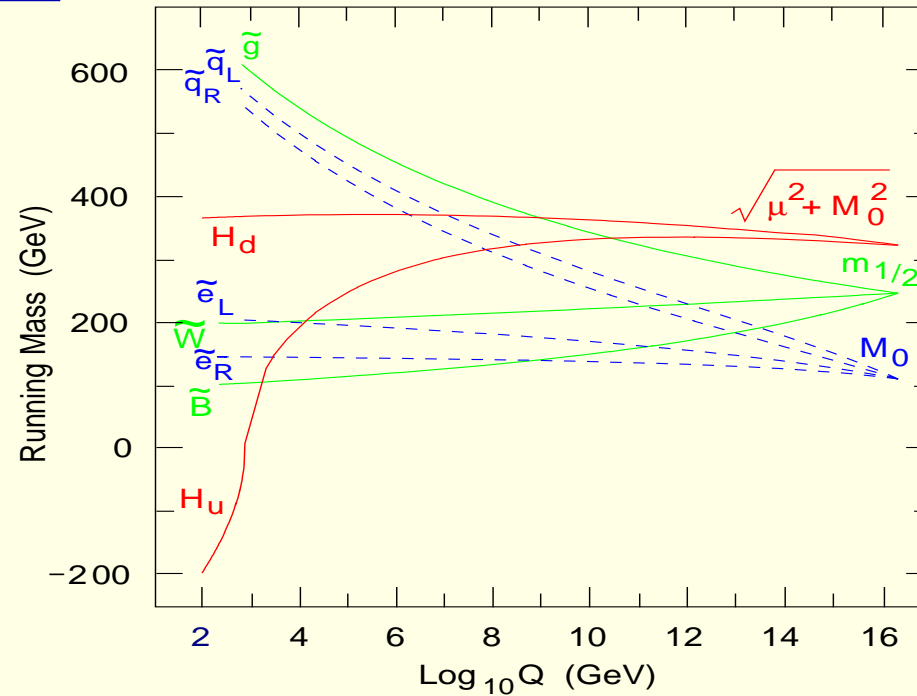


Figure 5: Evolution of the (soft) SUSY-breaking masses or masses-squared, showing how $m_{H_u}^2$ is driven < 0 at low $Q \sim \mathcal{O}(m_Z)$.

But, must one fine-tune the GUT scale parameters to get correct Z mass?

F measures the degree to which GUT parameters must be tuned. Want $F < 10$. This requires $m_{\tilde{t}} \lesssim 400$ GeV and a relatively light gluino.

For such $m_{\tilde{t}}$ SUSY predicts $m_h < 110$ GeV. This is a problem for

the MSSM for which the h is typically SM-like in its decays. To get $m_h > 114$ GeV requires $m_{\tilde{t}} > 800$ GeV and then $F > 50$.

- What is needed is a SUSY model for which the stop mass can be low but for which the resulting light $\lesssim 105$ GeV Higgs is not excluded by LEP.

LEP exclusion can be avoided by having unusual decays as seen earlier.

- **The NMSSM is perfect**

It is the h_1 that is light and SM-like and the a_1 is mainly singlet and has a small mass that is protected by a $U(1)_R$ symmetry. Large $B(h_1 \rightarrow a_1 a_1)$ is easy to achieve.

The many attractive features of the NMSSM are well known:

1. Solves μ problem: $W \ni \lambda \hat{S} \hat{H}_u \hat{H}_d \Rightarrow \mu_{\text{eff}} = \lambda \langle S \rangle$.
2. Preserves MSSM gauge coupling unification.

3. Preserves radiative EWSB.
4. Preserves dark matter (assuming R -parity is preserved).
5. Like any SUSY model, solves quadratic divergence hierarchy problem.
6. Has additional attractive features when $m_{h_1} \sim 90 - 105$ GeV is allowed because of $h_1 \rightarrow a_1 a_1$ decays with $m_{a_1} < 2m_b$:
 - (a) Allows minimal fine-tuning for getting m_Z (i.e. v) correct after evolving from GUT scale M_U . (R. Dermisek and J. F. Gunion, Phys. Rev. D 73, 111701 (2006) [arXiv:hep-ph/0510322])
 This is because \tilde{t}_1, \tilde{t}_2 can be light (~ 350 GeV is just right) . Also need $m_{\tilde{g}}$ not too far above 300 GeV.
 (In MSSM, such low stop masses are not acceptable since m_{h^0} would be below LEP limits; large $m_{\tilde{t}} \Rightarrow m_Z$ fine tuning would be large, especially if m_h is SM-like.)
 - (b) An a_1 with large $B(h_1 \rightarrow a_1 a_1)$ and $m_{a_1} < 2m_b$ can be achieved without fine-tuning of the A_λ and A_κ soft-SUSY breaking parameters

$(V \ni A_\lambda S H_u H_d + \frac{1}{3} A_\kappa S^3)$ that control the a_1 properties. (R. Dermisek and J. F. Gunion, Phys. Rev. D 75, 075019 (2007) [arXiv:hep-ph/0611142].)

The a_1 is largely singlet (e.g. 10% at amplitude level if $\tan \beta \sim 10$) and $\sim 7.5 \text{ GeV} \lesssim m_{a_1}$ (but below $2m_b$) in the best cases.

7. Of course, multi-singlet extensions of the NMSSM will expand the possibilities.

Indeed, typical string models predict a plethora of light a 's, light h 's and light $\tilde{\chi}$'s .

8. Many other non-Higgs decay modes of the h or h_1 have been proposed. Even sticking to SUSY, we have lots.

Models which preserve R -parity and thus dark matter possibility include:

(a) $h \rightarrow \tilde{\chi}_2^0 \tilde{\chi}_1^0$ followed by $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 f \bar{f}$ (S. Chang and T. Gregoire, arXiv:09030403): Turns out to be hard to accommodate given LEP constraints.

(b) $h \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \tilde{G} \tilde{G} \gamma \gamma \rightarrow \cancel{E}_T \gamma \gamma$: Can't recall others who have worked on this, but I consider it likely that LEP would have seen such decays for a light h in the mass range of interest for PEW perfection.

(c) $h \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \cancel{E}_T$: perfectly viable for non-unified gaugino masses, but LEP limit for invisibly decaying h is 114 GeV which is too heavy for PEW perfection.

Many other models also have dominant invisible h decay, but all suffer from the $m_h > 114$ GeV LEP limit for this mode that is less than ideal for PEW.

$\xi^2 = [\sigma(Zh)/\sigma(Zh_{\text{SM}})] \times B(h \rightarrow \cancel{E})$ curve for invisible mode is very steep, \Rightarrow allowed m_h does not decrease much until ξ^2 is quite small.

Models which violate R parity (and therefore require an alternative DM candidate than the $\tilde{\chi}_1^0$):

(a) There are too many to list systematically. A particularly nasty one is baryon-violating R -parity decays (L.M. Carpenter, D.E. Kaplan and E-J Rhee, arXiv:hep-ph/0607204) $h \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow (3j)(3j)$.

Such a multi-jet mode is least constrained by LEP ($m_h > 82$ GeV is the limit) and the lighter the h the better the agreement with precision data (especially dropping hadronic asymmetries).

Predictions regarding a light a and the NMSSM a_1

- Define the mass eigenstate: $a_1 = \cos \theta_A a_{MSSM} + \sin \theta_A a_S$.

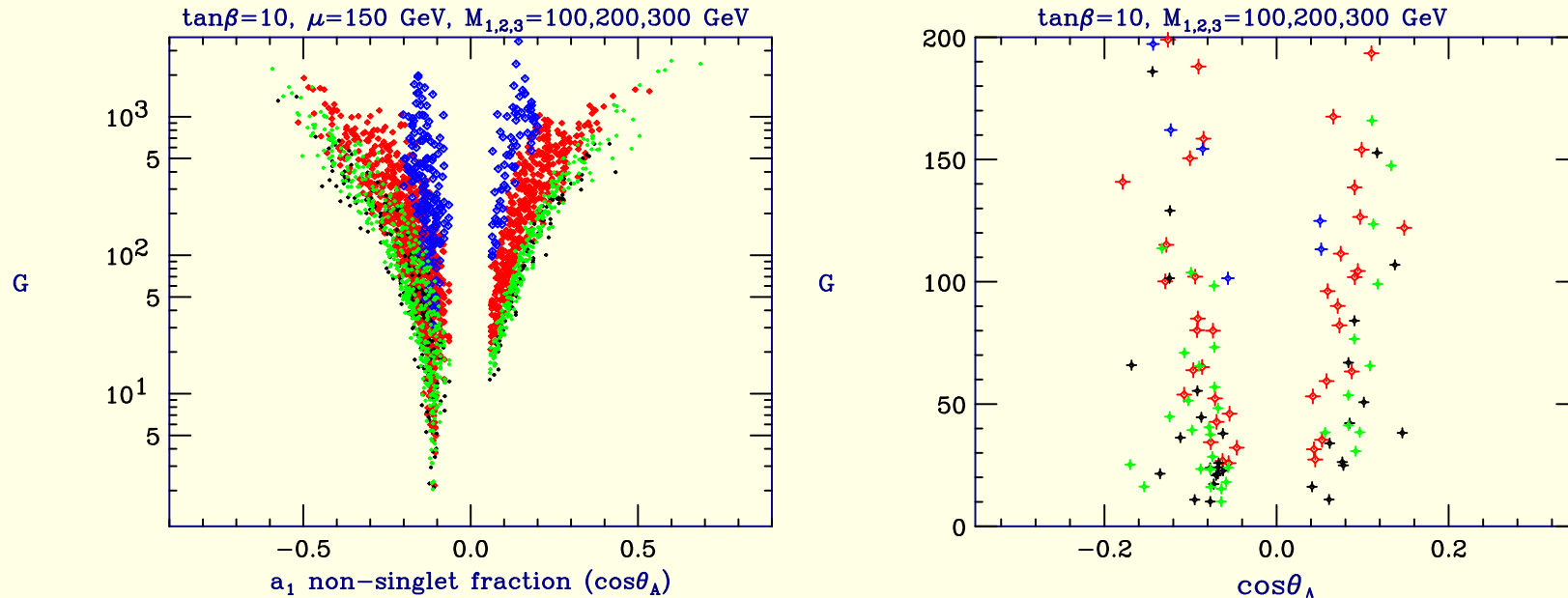


Figure 6: G vs. $\cos \theta_A$ for $M_{1,2,3} = 100, 200, 300$ GeV and $\tan \beta = 10$ from $\mu_{\text{eff}} = 150$ GeV scan (left) and for points with $F < 15$ (right) having $m_{a_1} < 2m_b$ and large enough $B(h_1 \rightarrow a_1 a_1)$ to escape LEP limits. The color coding is: blue = $m_{a_1} < 2m_\tau$; red = $2m_\tau < m_{a_1} < 7.5$ GeV; green = 7.5 GeV $< m_{a_1} < 8.8$ GeV; and black = 8.8 GeV $< m_{a_1} < 9.2$ GeV.

- In the figure, G is a measure (Dermisek+JFG: hep-ph/0611142) of the degree

to which A_λ and A_κ have to be fine tuned ("light- a_1 " fine tuning) in order to achieve required a_1 properties of $m_{a_1} < 2m_b$ and $B(h_1 \rightarrow a_1 a_1) > 0.7$. The plot of G vs. $\cos \theta_A$ shows a strong preference for $m_a > 7.5$ GeV and $\cos \theta_A \lesssim 0.1$ (for $\tan \beta = 10$). **Note the strict lower bound on $\cos \theta_A$ needed for $B(h_1 \rightarrow a_1 a_1) > 0.7$.**

- Define a generic coupling to fermions by

$$\mathcal{L}_{aff} \equiv iC_{aff} \frac{ig_2 m_{f^-}}{2m_W} \bar{f} \gamma_5 f a, \quad \text{then} \quad C_{abb} = \cos \theta_A \tan \beta \quad (1)$$

At large $\tan \beta$, SUSY corrections $C_{abb} = C_{abb}^{tree} [1/(1 + \Delta_b^{SUSY})]$ can be large and either suppress or enhance C_{abb} relative to $C_{a\tau-\tau+}$. Will ignore.

- The extracted C_{abb} limits (JFG, arXiv:0808.2509 and JFG+Dermisek, in preparation; see also Ellwanger and Domingo, arXiv:0810.4736) appear in Fig. 7.
- The most unconstrained region is that with $m_a > 8$ GeV, especially $9 \text{ GeV} < m_a < 12 \text{ GeV}$.

The $9 \text{ GeV} m_a < 2m_B$ portion of the latter is the same as the region with least "light- a_1 " fine-tuning in the NMSSM.

- One needs to achieve limits of $C_{abb\bar{b}} < 0.3$ to rule out the a_1 of the $C_{abb\bar{b}} = \cos \theta_A \tan \beta \lesssim 1$ (a number which applies for $\tan \beta > 3$) scenarios preferred to achieve small light- a_1 finetuning.

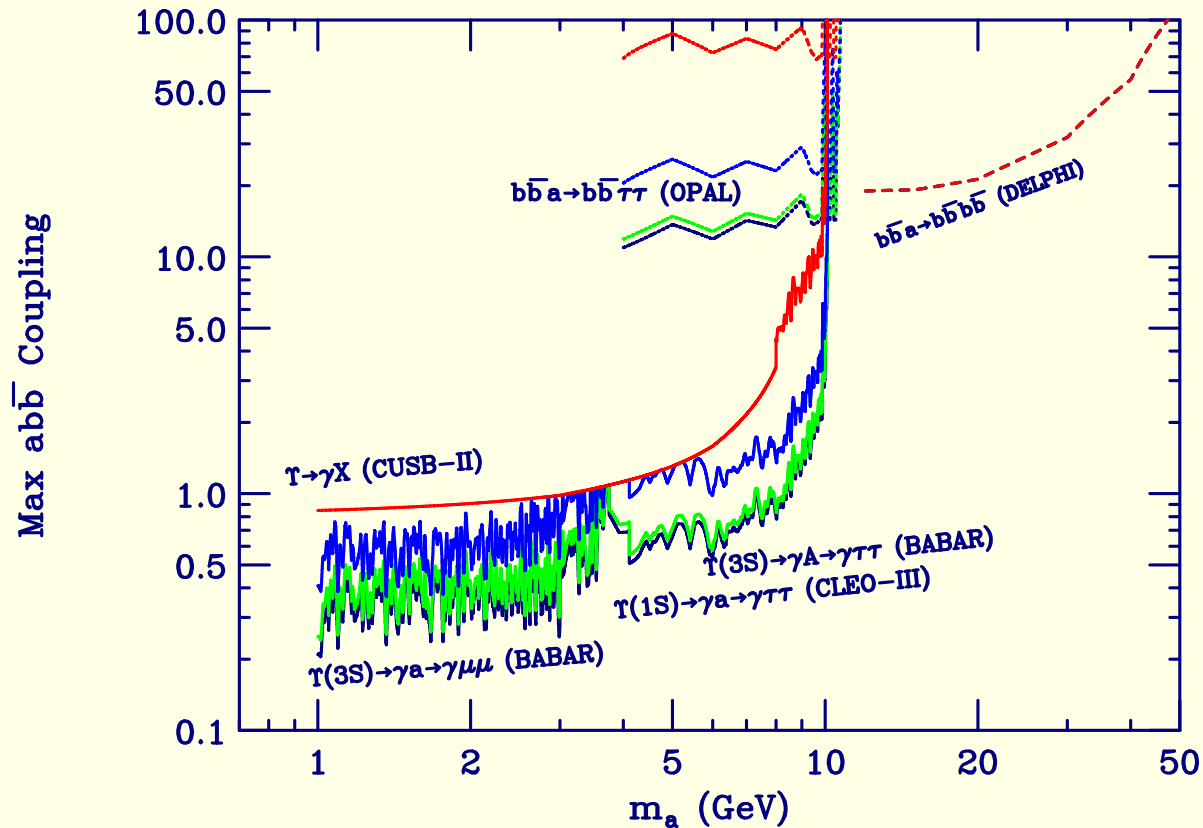


Figure 7: Limits on $C_{abb\bar{b}}$ from JFG, arXiv:0808.2509 and JFG+Dermisek, in preparation. These limits include recent BaBar $\Upsilon_{3S} \rightarrow \gamma \mu^+ \mu^-$ and $\gamma \tau^+ \tau^-$ limits. Color code: $\tan \beta = 0.5$; $\tan \beta = 1$; $\tan \beta = 2$; $\tan \beta \geq 3$.

- In the $\sim 9 \text{ GeV} \lesssim m_a \lesssim 12 \text{ GeV}$ region only the OPAL limits are relevant.

Those presented depend upon how the $a \leftrightarrow \eta_b$ states mixing is modeled. A particular model (Drees+Hikasa: Phys.Rev.D41:1547,1990) is employed.

Perhaps now that the first η_b state has been observed, this region can be better pinned down. I have not incorporated recent work by Domingo *et al.* (arXiv:0810.4736) which models this mixing in a manner consistent with the available information. In any case, models predict many η -type states in this region, not just the one that has been observed.

- Given $C_{abb\bar{}}$ limits, an interesting question is whether there is any possibility that a light a could be responsible for the observed a_μ discrepancy which is of order $\Delta a_\mu \sim 30 \times 10^{-10}$.

For this, large $C_{abb\bar{}}$ is needed.

The plotted limits (mainly BaBar at up near $m_a \sim 9 \text{ GeV}$) suggest that it is generically possible from $C_{abb\bar{}}$ limits if $m_a > 9 \text{ GeV}$, but is not possible in the NMSSM scenarios with small light- a_1 fine-tuning since they do not have large $C_{abb\bar{}}$.

- We will see that $B(a_1 \rightarrow \mu^+ \mu^-)$ is an interesting quantity.

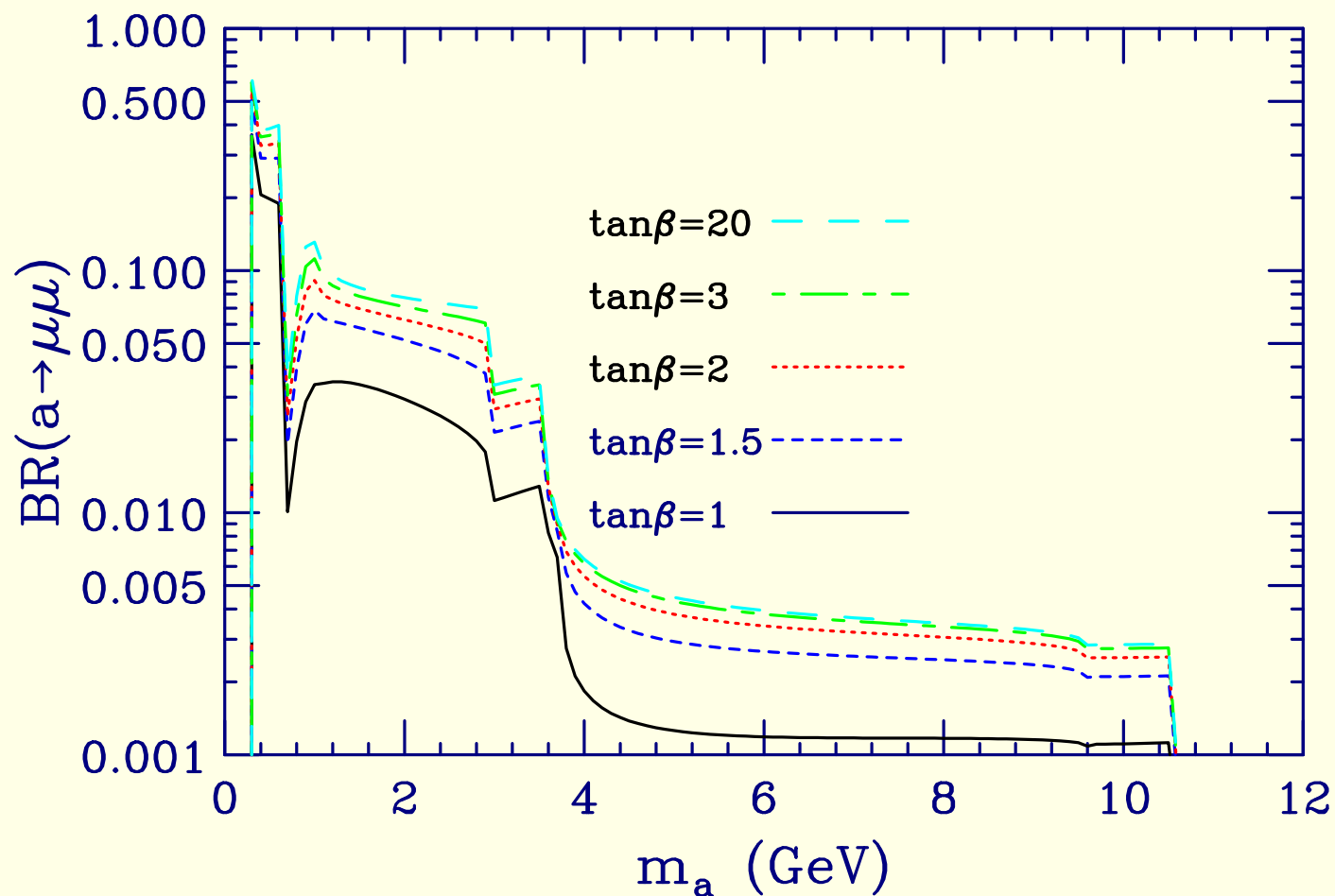


Figure 8: $B(a \rightarrow \mu^+ \mu^-)$ for various $\tan\beta$ values.

- It will also become important to know about $B(a_1 \rightarrow \tau^+\tau^-)$. Note values at high $\tan\beta$ of ~ 0.75 for $m_a \gtrsim 10$.

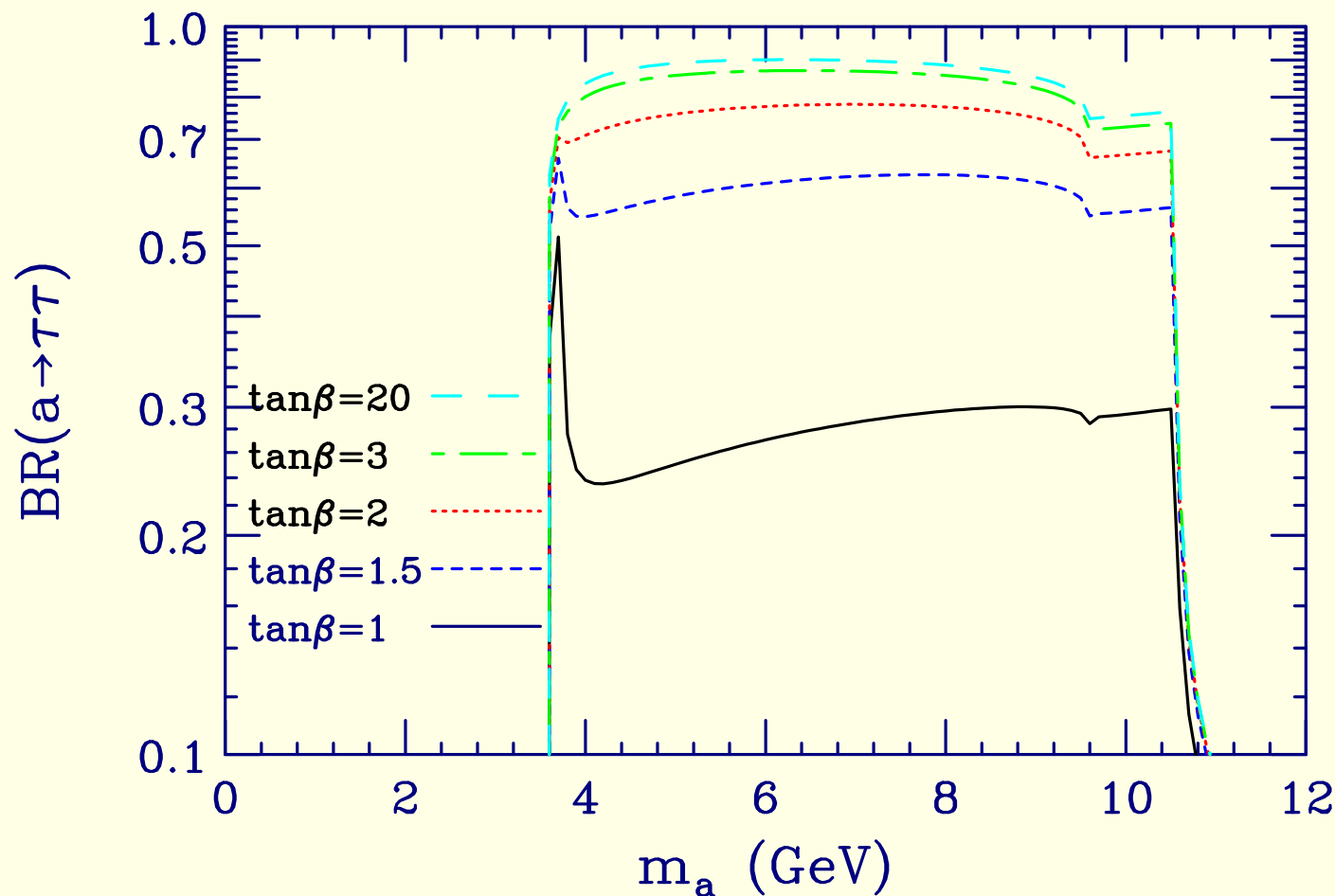


Figure 9: $B(a \rightarrow \tau^+\tau^-)$ for various $\tan\beta$ values.

- More on the strong BaBar limits on $B(\Upsilon_{3S} \rightarrow a\gamma)B(a \rightarrow \mu^+\mu^-)$ that become very constraining for $m_a < 2m_\tau$.

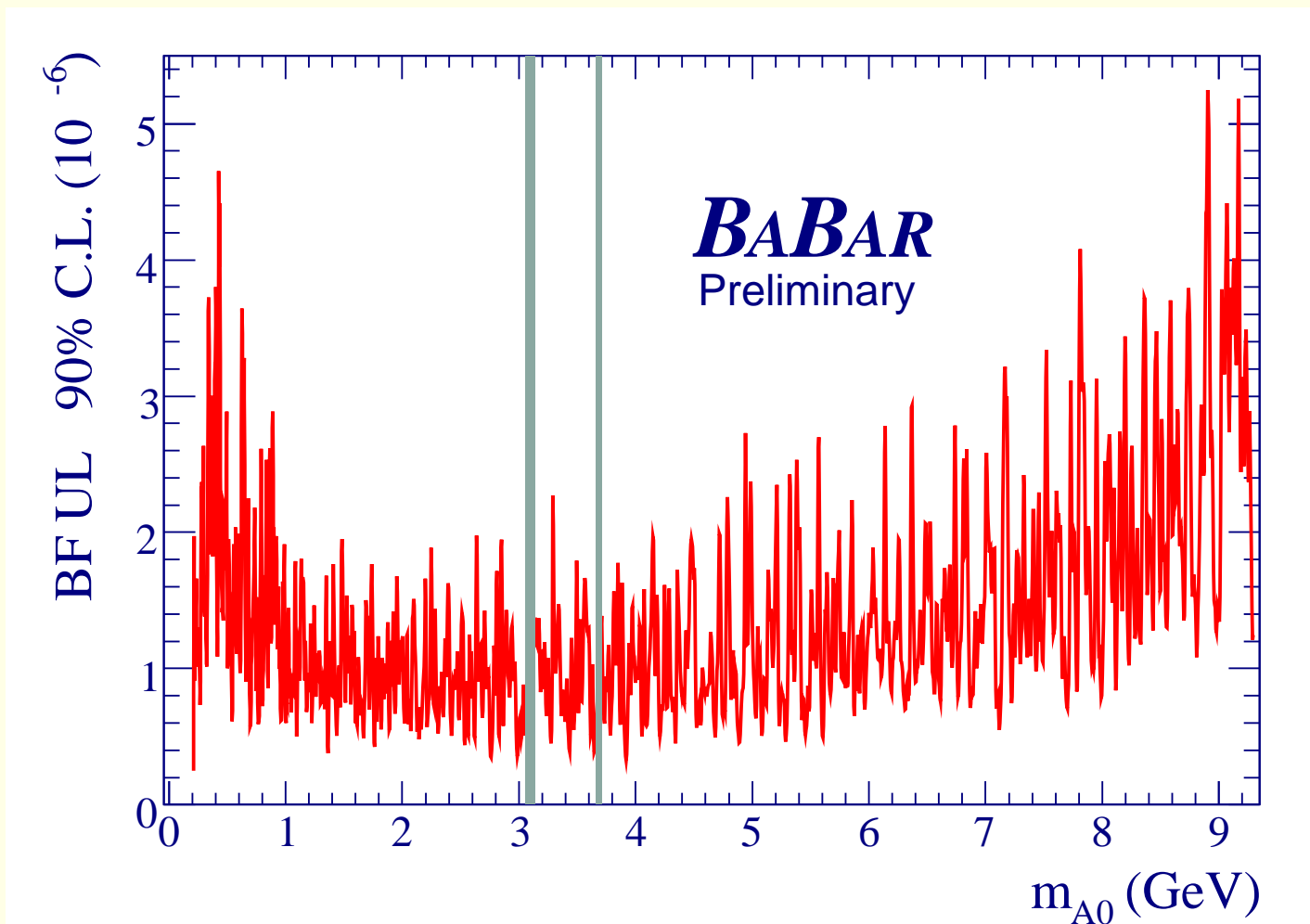


Figure 10: BaBar limits on $B(\Upsilon_{3S} \rightarrow \gamma a)B(a \rightarrow \mu^+\mu^-)$.

For $m_a < 2m_\tau$, the limits are below 2×10^{-6} except for very low m_a .

A comparison to NMSSM predictions \Rightarrow most NMSSM scenarios with $B(h_1 \rightarrow a_1 a_1) > 0.7$ and $m_{a_1} < 2m_\tau$ are eliminated; only a few at $\tan\beta \lesssim 3$ survive.

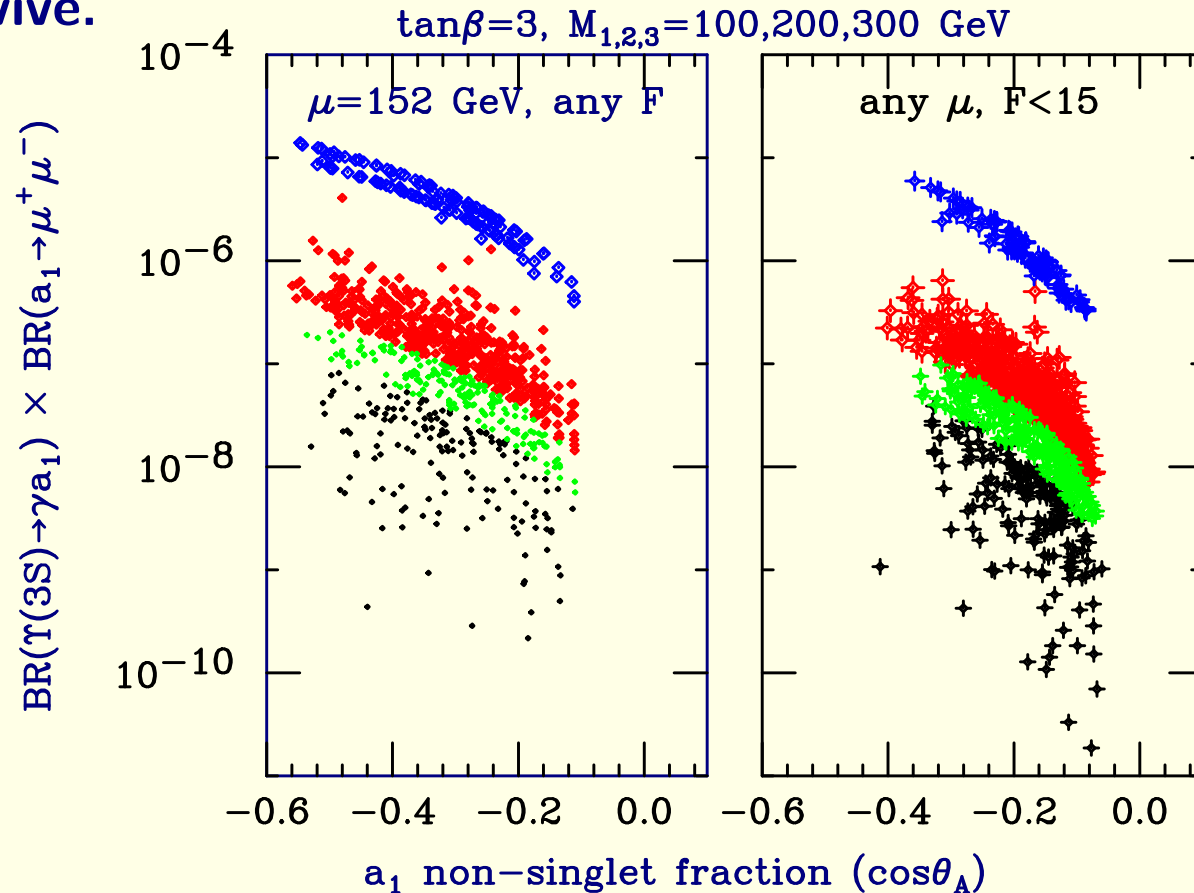


Figure 11: For $\tan\beta = 3$, we plot $B(\Upsilon_{3S} \rightarrow \gamma a_1) \times B(a_1 \rightarrow \mu^+ \mu^-)$ for NMSSM scenarios with various ranges for m_{a_1} . Color code: $m_{a_1} < 2m_\tau$; $2m_\tau < m_{a_1} < 7.5$ GeV; 7.5 GeV $< m_{a_1} < 8.8$ GeV; 8.8 GeV $< m_{a_1} < 2m_B$ GeV. The left plot comes from an A_λ, A_κ scan holding $\mu_{eff}(m_Z) = 152$ GeV fixed. The right plot shows results for $F < 15$ scenarios with $m_{a_1} < 2m_B$.

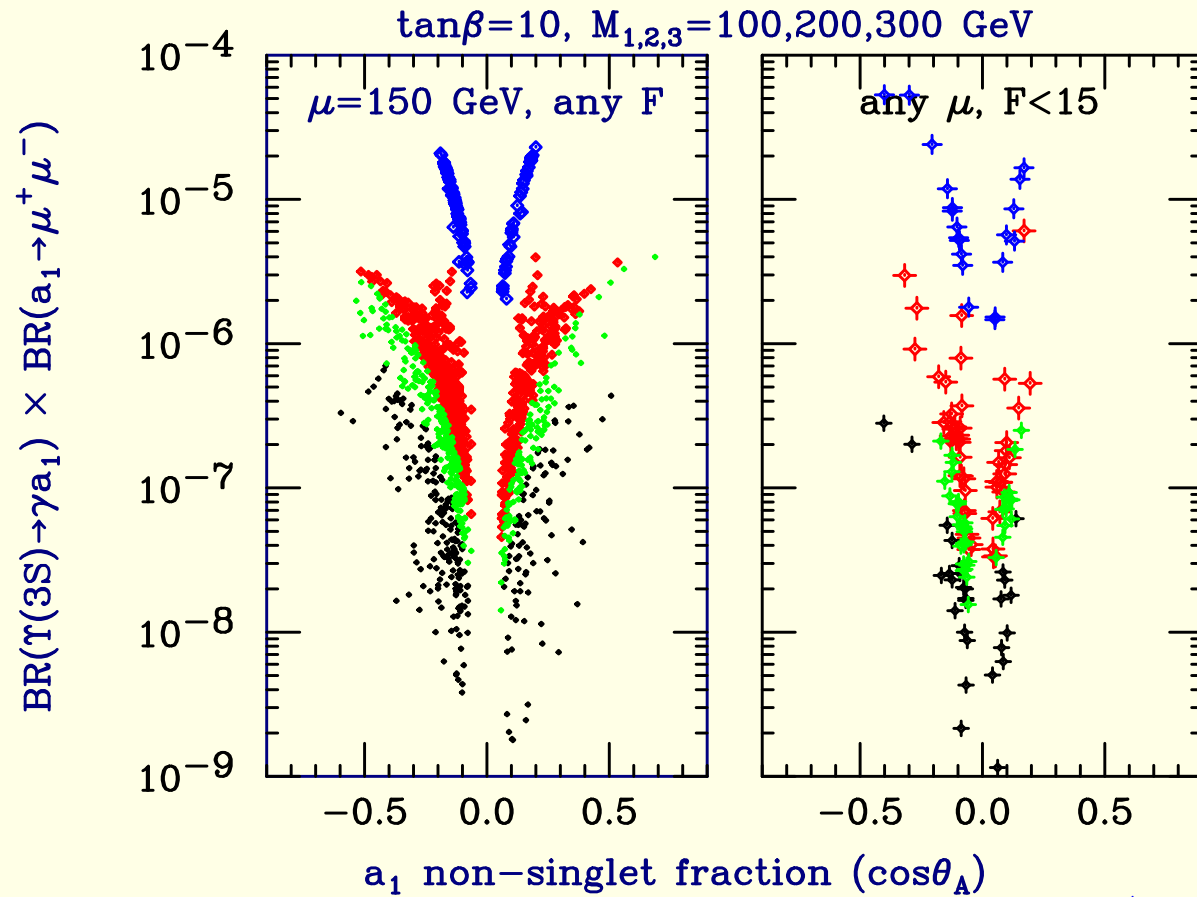


Figure 12: For $\tan\beta = 10$ we plot $B(\Upsilon_{3S} \rightarrow \gamma a_1) \times B(a_1 \rightarrow \mu^+ \mu^-)$ for NMSSM scenarios with various ranges for m_{a_1} . Color code: $m_{a_1} < 2m_\tau$; $2m_\tau < m_{a_1} < 7.5 \text{ GeV}$; $7.5 \text{ GeV} < m_{a_1} < 8.8 \text{ GeV}$; $8.8 \text{ GeV} < m_{a_1} < 2m_B \text{ GeV}$. The left plot comes from an A_λ, A_κ scan holding $\mu_{eff}(m_Z) = 150 \text{ GeV}$ fixed. The right plot shows results for $F < 15$ scenarios with $m_{a_1} < 2m_B$.

- Thus, we have a convergence whereby low “light- a ” fine tuning in the NMSSM and direct $\Upsilon_{3S} \rightarrow \gamma\mu^+\mu^-$ limits single out the $m_a > 2m_\tau$ part of parameter space.

LHC studies of light h NMSSM scenarios should (and have) focused on this case.

With regard to the a itself, we should focus on Tevatron and LHC probes of a light a with $2m_\tau < m_a < 2m_B$.

This is not to say that the Tevatron and LHC cannot be sensitive to $m_a < 2m_\tau$:

1. $B(a \rightarrow \mu^+\mu^-)$ is much larger. **BUT**
2. Acceptance is presumably considerably smaller because of p_T distributions for the μ 's shifting down.
3. Backgrounds are presumably larger.

Studies of $m_a < 2m_\tau$ cases at hadron colliders are worth pursuing since they might completely eliminate all such NMSSM ideal Higgs scenarios.

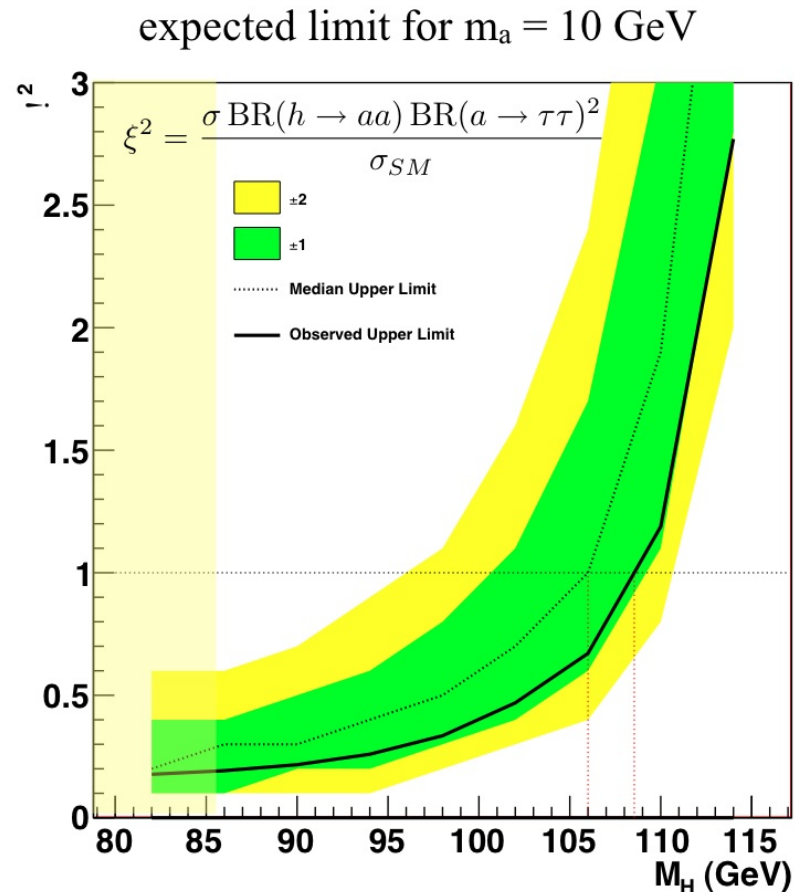
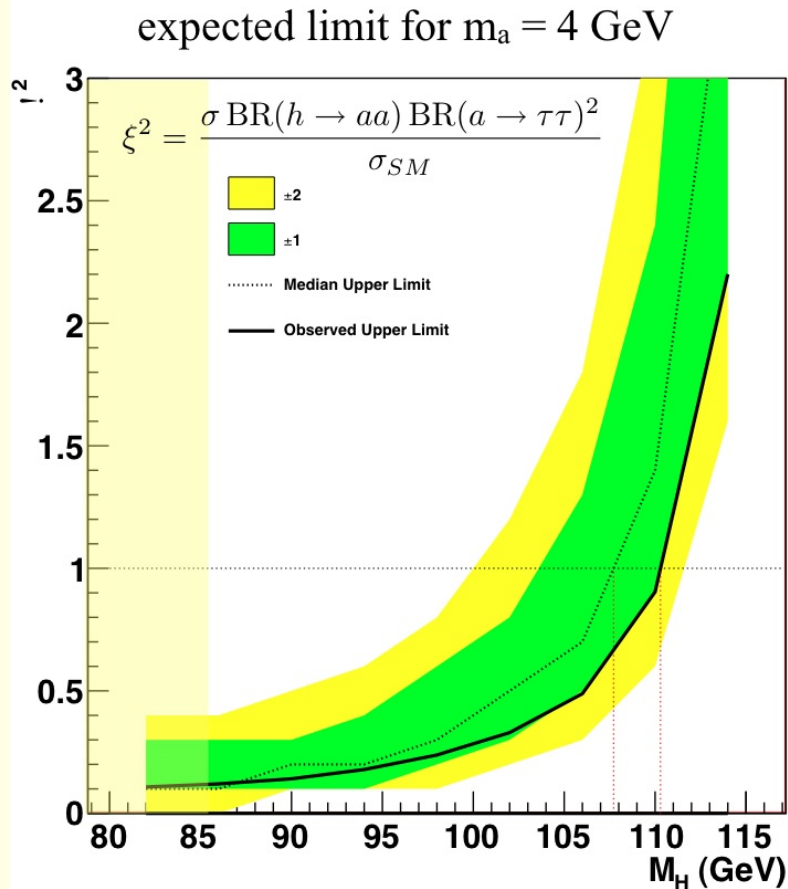
Here we will focus on $m_a > 2m_\tau$.

- In fact, results from ALEPH that came out last week (Kyle Cranmer, Nov. 3 seminar) further shift the focus to high m_a in the NMSSM context.

Expected limits @ $m_a = 4, 10$ GeV

Seeing no sign of excess, we proceed to set limits

- ! Here, we make reference to background acceptance uncertainties in MSSM Higgs analysis. (Statistical errors dominate, systematics make little difference in result)



- Comparison to NMSSM ideal scenarios:

1. $m_h \sim 95 \text{ GeV} - 103 \text{ GeV}$ to minimize electroweak m_Z finetuning.
2. Large enough $B(h \rightarrow b\bar{b}) \sim 0.15 - 0.2$ to explain 2.3σ LEP excess.
3. $9 \lesssim m_a \lesssim 2m_B$ to fully minimize light- a finetuning.

In this case, we typically have

1. $\sigma(h)/\sigma(h_{\text{SM}}) \sim 0.92 - 1.0$
2. $B(h \rightarrow aa) \sim 0.8 - 0.85$
3. $B(a \rightarrow 2\tau) \sim 0.75 - 0.8$

Together, these yield $\xi^2 \sim 0.43 - 0.55$.

- Thus, while $m_a \sim 4 \text{ GeV}$ is still ok if $m_h \sim 105 \text{ GeV}$ and ξ^2 is in the above range, **ALEPH limits tend to push into the higher m_a part of model space that is most ideal with respect to light- a finetuning perspective and explaining the LEP 2.3σ excess near 100 GeV.**

If all LEP experiments perform this kind of analysis and combine results will they rule out this corner?

Hadron collider constraints on a light a

- As we have seen, the Upsilon constraints on a light a run out for $m_a > M_{\Upsilon_{3S}} - \delta$. This leaves open the possibility that Δa_μ could be explained by a light a if $C_{abb\bar{b}}$ is big in this region. Remarkably, existing Tevatron data rule out this possibility (JFG+Dermisek, in preparation). And LHC constraints on the a or a_1 are likely to be even stronger.
- At a hadron collider, one studies $\mu^+\mu^-$ pair production and tries to reduce the heavy flavor background by isolation cuts on the muons. Various studies of Υ production have been performed and CDF has even done an analysis in which they look for a very narrow ϵ (a hypothesized particle of a non-SUSY model) over the region $6.3 < m_\epsilon < 9$ GeV. The latest CDF limits from $L = 630 \text{ pb}^{-1}$ of data on $R \equiv \sigma(\epsilon)B(\epsilon \rightarrow \mu^+\mu^-)/\sigma(\Upsilon_{1S})B(\Upsilon_{1S} \rightarrow \mu^+\mu^-)$ rule out the old peak at $m_\epsilon = 7.2$ GeV and can be adopted to limit this same ratio for a general a or the NMSSM a_1 .

- **Ingredients:**

- First, we need the cross sections. These are basically from gg fusion with gga coupling induced by quark loops. Higher order corrections, both virtual and real (*e.g.* for the latter $gg \rightarrow ag$) are, however, quite significant.

Main points are:

1. Isolation cuts on μ 's do not seem to exclude NLO real radiation diagrams (based on CDF, ATLAS, CMS Υ efficiencies and fact that $\sigma(\Upsilon)$ has many components involving one or more extra final state g or q).
 2. Slow energy variation. At $m_a = 10$ GeV and $\tan\beta = 10$, one finds $\sigma_{NLO}(1.96, 7, 10, 14 \text{ TeV}) \sim 1.5 \times 10^5, 5 \times 10^5, 7 \times 10^5, 9 \times 10^5$ pb.
 3. For NMSSM, multiply by $(\cos\theta_A)^2$.
- Then, we must know $B(a \rightarrow \mu^+\mu^-)$, which we plotted earlier, a rough value being 0.003 for $m_a > 2m_\tau$ and $\tan\beta > 2$.
 - We need efficiencies for detecting the μ^+ and μ^- at given m_a .
 - We must know the background, which mainly derives from heavy flavor production, especially $b\bar{b}$ where the b 's decay semi-leptonically.

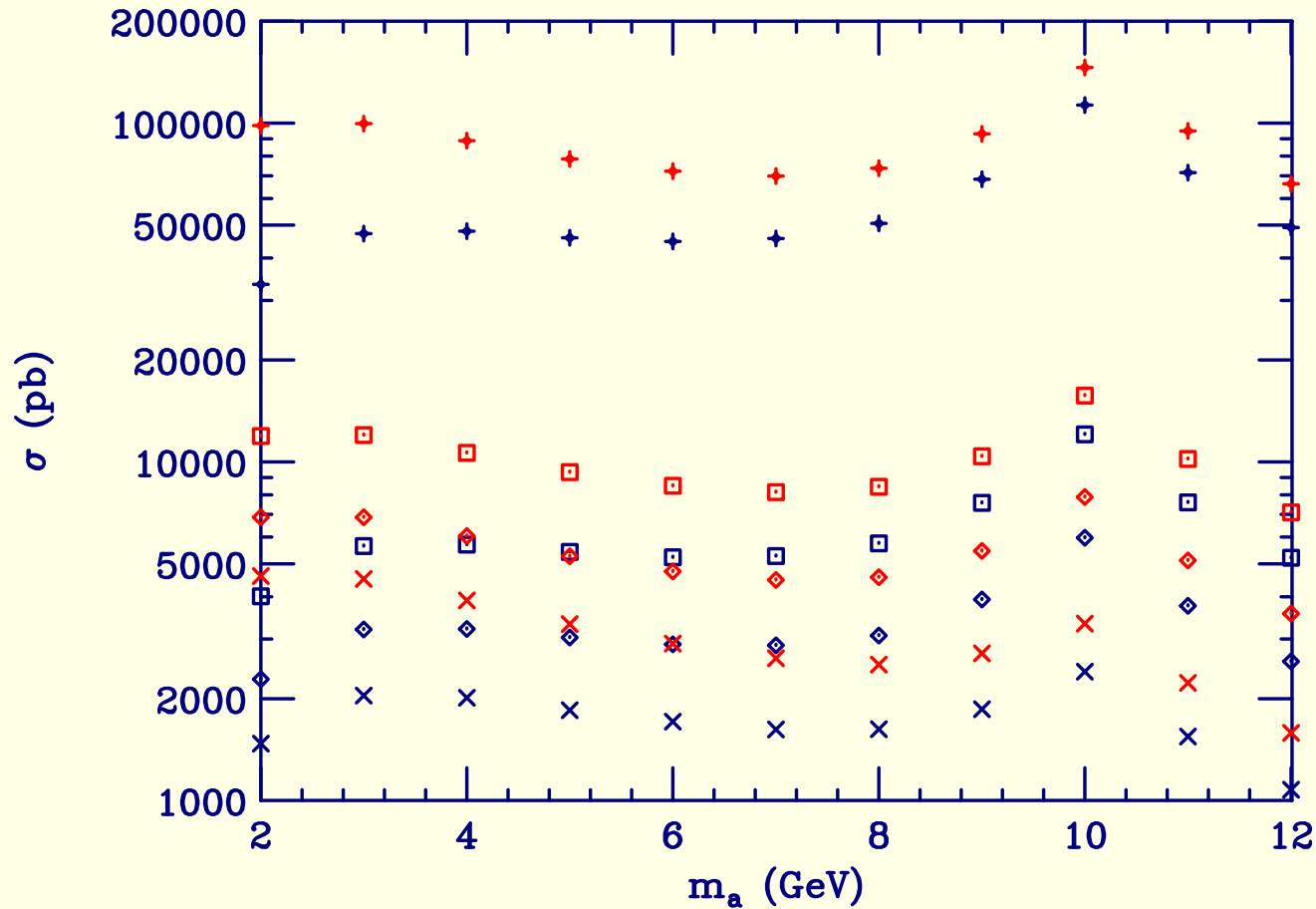


Figure 13: Tevatron cross sections for $\tan\beta = 1, 2, 3, 10$ (lowest to highest point sets). For each m_a and $\tan\beta$ value, the lower (higher) point is the cross section without (with) resolvable parton final state contributions.

For later reference when we discuss LHC:

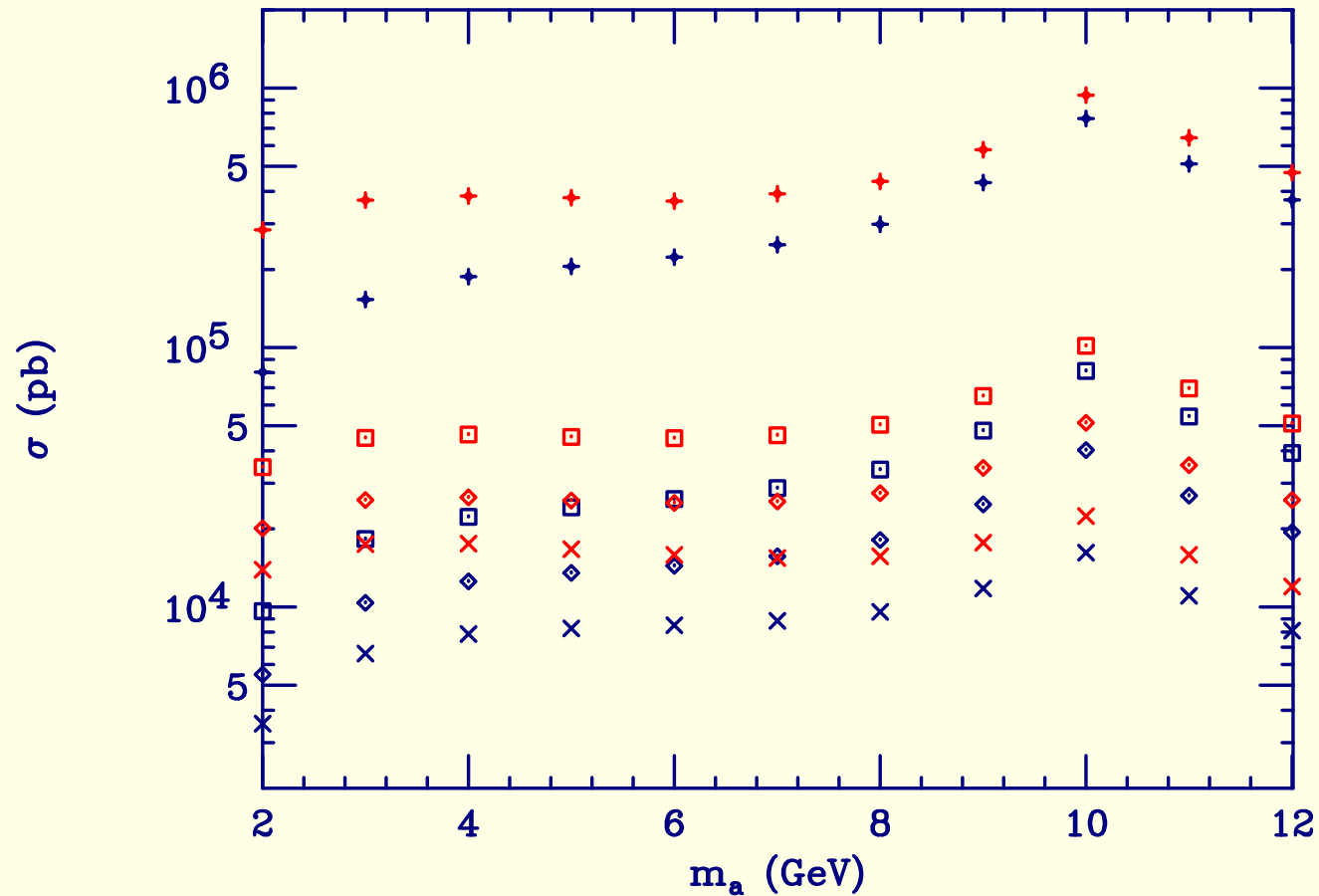


Figure 14: LHC, $\sqrt{s} = 10$ TeV cross sections for $\tan\beta = 1, 2, 3, 10$ (lowest to highest point sets). Factor of about $7 \times$ Tevatron at higher m_a .

Putting it all together gives:

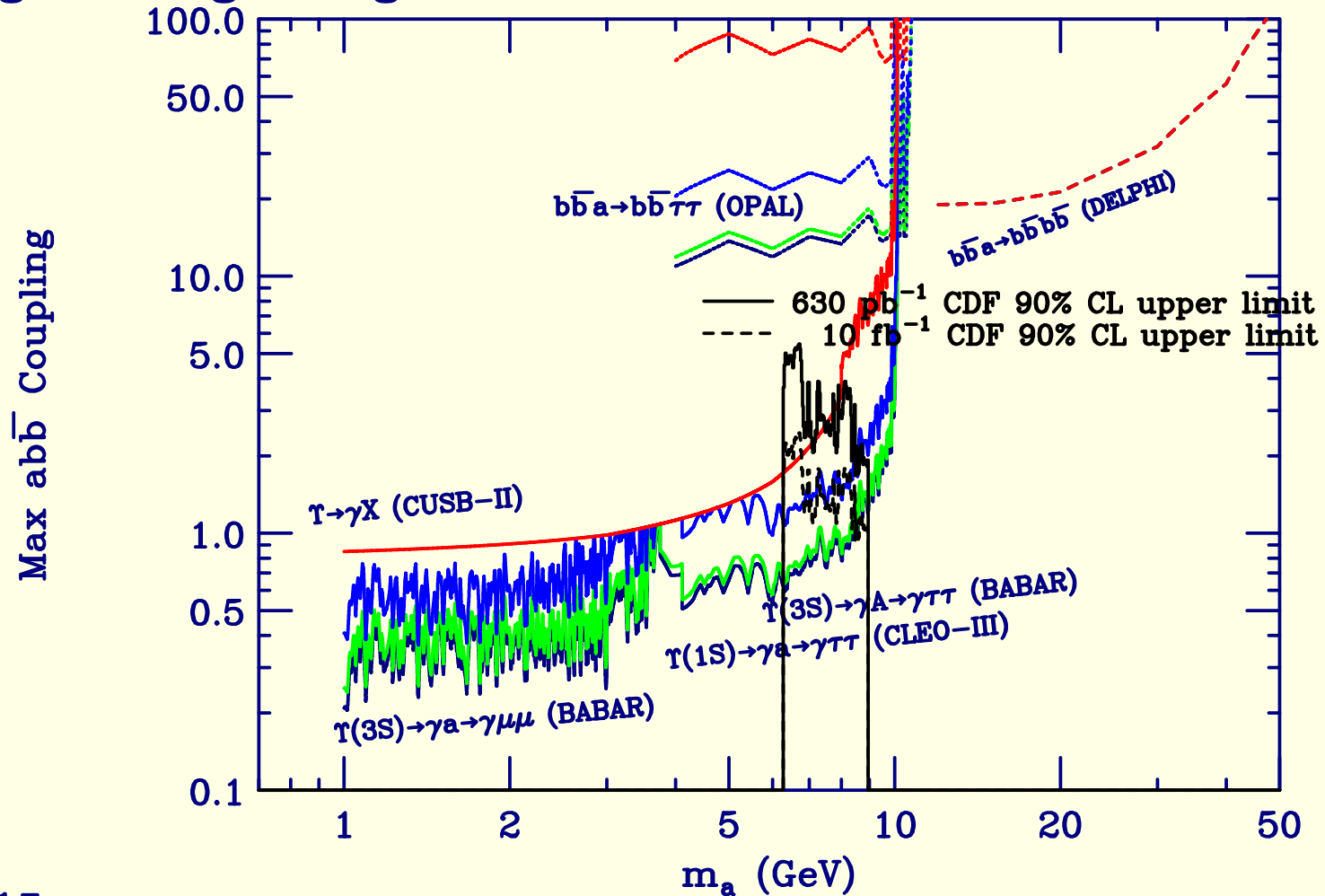


Figure 15: Tevatron limits compared to previous plot limits for $\tan \beta = 0.5, 1, 2, \geq 3$. Tevatron at $L = 10 \text{ fb}^{-1}$ competes with BaBar for $m_a \sim 9 \text{ GeV}$ and would win above that. Indeed, **The $L = 10 \text{ fb}^{-1}$ statistically extrapolated limits are approaching the $C_{abb} = \tan \beta \cos \theta_A \sim 1$ level that impacts the most preferred NMSSM scenarios.**

For $M_{\mu+\mu^-} > 9$ GeV, CDF did not perform the R analysis. Instead, we use the event number plots that extend to larger $M_{\mu+\mu^-}$. We ask for the $|C_{abb\bar{b}}|$ limits assuming no 90% CL (1.686σ) fluctuation in S/\sqrt{B} -optimized m_a interval of $2\sqrt{2}\sigma_r$, where σ_r is the $M_{\mu+\mu^-}$ resolution.

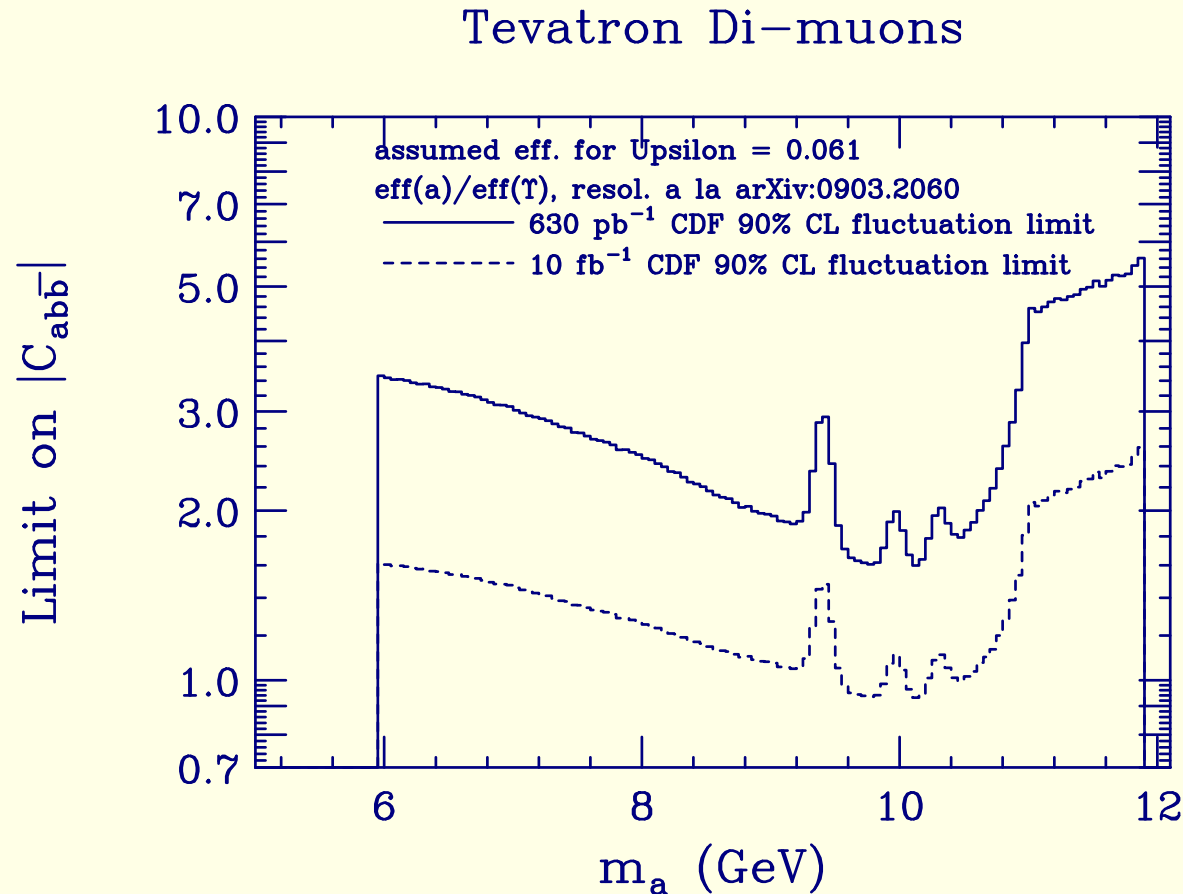


Figure 16: $L = 630 \text{ pb}^{-1}$ and 10 fb^{-1} limits based on no 1.686σ excess in optimal interval.

We see that in the region below 12 GeV where a light a might have explained Δa_μ if $C_{abb\bar{b}} \gtrsim 32$, current Tevatron data forbids such a large $C_{abb\bar{b}}$. One can finally conclude that Δa_μ cannot be due to a light a .

What about the LHC? There have been studies by CMS and ATLAS, and for reasons that I am still trying to explore with the experimentalists the di-muon background in the CMS studies is larger than that in the ATLAS studies. Also, only ATLAS has presented public results — see Fig. 17.

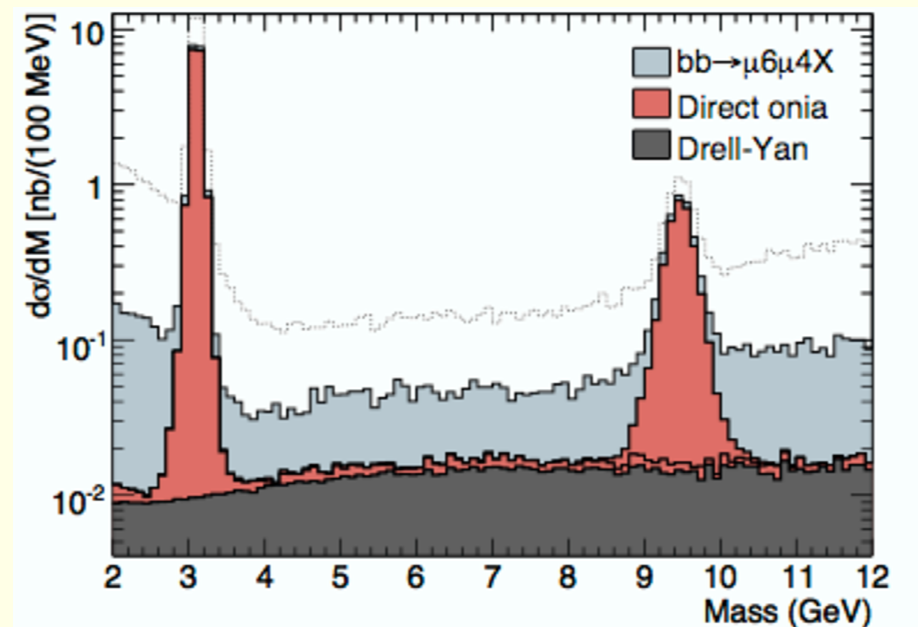


Figure 17: ATLAS dimuon spectrum prediction after corrections for acceptance and efficiencies (D. D. Price, arXiv:0808.3367 [hep-ex].).

Consider $\tan\beta = 10$ and $\cos\theta_A = 0.1$ (middle range of most preferred NMSSM models).

- After accounting for efficiency \times tracking factor of $\sim 50\%$ (vs. CDF 6%), the need to double plotted continuum background which was only from $b\bar{b}$ (in particular, did not include $c\bar{c}$), and the resolutions $\sigma_r(M_{\mu^+\mu^-})$ (54 MeV at J/ψ and 170 MeV at Υ_{1S}), we compute the number, $N_{\Delta M_{\mu^+\mu^-}}$, of events in an interval of total width $\Delta M_{\mu^+\mu^-} = 2\sqrt{2}\sigma_r$ (the interval that maximizes S/\sqrt{B}).

We obtain background levels of 2121, 23519, and 4819 at 8 GeV, $M_{\Upsilon_{1S}}$ and 10.5 GeV, respectively. Note: at Υ_{1S} peak can use $\Upsilon_{1S} \rightarrow e^+e^-$ to independently measure this background.

We compute $\sqrt{N_{\Delta M_{\mu^+\mu^-}}}$ 1σ errors of 45, 153 and 69 at 8 GeV, $M_{\Upsilon_{1S}}$ and 10.5 GeV, respectively.

- We now consider the $a \rightarrow \mu^+\mu^-$ signal rates.

From Fig. 14, we see that at $\tan\beta = 10$ the total a cross section ranges

from about $4.2 \times 10^5 \text{ pb}(\cos \theta_A)^2 \sim 4200 \text{ pb}$ at $m_a = 8 \text{ GeV}$ to $\sim 8500 \text{ pb}$ at $m_a \lesssim 2m_B$ for $\sqrt{s} = 14 \text{ TeV}$. Including $B(a \rightarrow \mu^+\mu^-) \sim 0.003$ we get $\sigma(gg \rightarrow a \rightarrow \mu^+\mu^-) \sim 12 - 25 \text{ pb}$ in $m_a \in [8 \text{ GeV} - 2m_B]$

Multiplying by the $Erf(1) = 0.8427$ acceptance factor for the ideal interval being employed and using $L = 10 \text{ pb}^{-1}$, we obtain a event numbers of 101, 185 and 211 at $m_a = 8 \text{ GeV}$, $M_{\Upsilon_{1S}}$ and 10.5 GeV , respectively.

The statistical significances of the a peaks for $L = 10 \text{ pb}^{-1}$ are then 2.2σ , 1.2σ and 3.0σ , respectively. But, small S/B values esp. at Υ_{1S} peak.

- Of course, we currently expect that substantial early running will mostly take place at $\sqrt{s} = 7 \text{ TeV}$ and $\sqrt{s} = 10 \text{ TeV}$.

As noted earlier, lower \sqrt{s} implies a somewhat smaller a cross section in the $[8 \text{ GeV}, 2m_B]$ mass interval on which we are focusing. Roughly, relative to $\sqrt{s} = 14 \text{ TeV}$, the a cross section decreases by a factor of ~ 1.3 at $\sqrt{s} = 10 \text{ TeV}$ and a factor of ~ 1.7 at $\sqrt{s} = 7 \text{ TeV}$ in this mass interval.

Since the backgrounds are also basically gg fusion induced, we presume that these same factors will apply to them. At $\sqrt{s} = 10 \text{ TeV}$ ($\sqrt{s} = 7 \text{ TeV}$)

this then will reduce the statistical significances given above by a factor of $1/\sqrt{1.3}$ ($1/\sqrt{1.7}$).

The statistical significances at $m_a = 8$ GeV, $M_{\Upsilon_{1S}}$ and 10.5 GeV are, respectively, then 2.0σ , 1.1σ , 2.7σ at 10 TeV and 1.7σ , 0.9σ , 2.3σ at 7 TeV.

At 10 TeV (7 TeV), to reach the 5σ signal level for $\tan\beta \cos\theta_A = 1$ at $m_a = M_{\Upsilon_{1S}}$ would require *only* $L = 207$ pb $^{-1}$ ($L = 309$ pb $^{-1}$).

Such integrated luminosities are quite likely to be achieved after a year or two of LHC early operation.

- We are somewhat surprised by the small integrated luminosities that we estimate for 5σ effects at ATLAS.

For example, at $m_a = 10.5$ GeV, about $L = 90$ fb $^{-1}$ is needed at Tevatron energies vs. only 34 pb $^{-1}$ required at 10 TeV at ATLAS.

Obvious factors causing increased or decreased significance at ATLAS include: ~ 5 in increase in gg induced xsecs due to larger \sqrt{s} ; factor of ~ 8

increase in detection efficiency (50% at ATLAS vs. 6% at CDF); factor of ~ 3 decrease due to worse ATLAS resolution (~ 170 MeV at ATLAS vs. ~ 52 MeV at CDF). Net factor of ~ 13 would imply that 90 fb^{-1} at CDF would be equivalent to $\sim 6.6 \text{ fb}^{-1}$ at ATLAS. **But this is still a factor of about 180 more than the $L = 34 \text{ pb}^{-1}$ estimate above.**

- One can repeat this kind of analysis using CMS inputs. We obtain the following comparisons.

Table 2: Comparison of statistical significances for $C_{abb\bar{b}} = \cos \theta_A \tan \beta = 1$

Case	$m_a = 8 \text{ GeV}$	$m_a = M_{\Upsilon_{1S}}$	$m_a \lesssim 2m_B$
Tevatron, $L = 10 \text{ fb}^{-1}$	0.9	0.7	1.7
CMS, LHC7, $L = 10 \text{ pb}^{-1}$	0.27	0.18	0.57
CMS, LHC10, $L = 10 \text{ pb}^{-1}$	0.31	0.21	0.65
CMS, LHC14, $L = 10 \text{ pb}^{-1}$	0.36	0.24	0.75
ATLAS LHC7, $L = 10 \text{ pb}^{-1}$	1.7	0.9	2.3
ATLAS LHC10, $L = 10 \text{ pb}^{-1}$	2.0	1.1	2.7
ATLAS LHC14, $L = 10 \text{ pb}^{-1}$	2.2	1.2	3.0

$C_{abb\bar{b}} \sim 0.2$ requires $[(1/0.2)^2]^2 \sim 625 \times$ more L to reach same levels.

NMSSM models in which several, perhaps many, Higgses carry the ZZ coupling

These arise for $\tan\beta < 3$. (R. Dermisek and J. F. Gunion, arXiv:0811.3537 [hep-ph].)

- It is possible to have h_1, h_2, h^+ all light but escaping LEP and Tevatron detection by virtue of decays to a_1 with $m_{a_1} < 2m_b$.
- h_1 need not be exactly SM-like — h_2 can be light enough (~ 100 GeV) for precision electroweak when $g_{h_2 WW}^2$ is substantial.
- Relevant scenarios often arise for $C_{abb\bar{b}} \gtrsim 1$, especially if $\tan\beta = 2$. Current limits imply that $m_{a_1} \gtrsim 9$ GeV is needed for $C_{abb\bar{b}} \sim 2$ to be ok. However, low $\tan\beta$ scenarios also arise for very small $C_{abb\bar{b}} \sim 0.2$, for which exclusion via direct a searches is very hard.
- **The multiple LEP (and Tevatron) escapes:**
 1. $B(h_1 \rightarrow a_1 a_1)$ is large, and $e^+e^- \rightarrow Zh_1 \rightarrow Za_1 a_1 \rightarrow Z4\tau$ is only constrained for $m_{4\tau} < 85$ GeV (recall decreased $B(a \rightarrow 2\tau)$ at low $\tan\beta$). Limit is lower if ZZh_1 coupling is somewhat suppressed.

2. $B(h^+ \rightarrow W^+ a_1)$ is often large, and $e^+ e^- \rightarrow h^+ h^- \rightarrow W^+ W^- a_1 a_1$ with $a_1 \rightarrow 2\tau$ was not directly searched for.
3. $B(h^+ \rightarrow \tau^+ \nu)$ is often significant (but never dominant) and for cases with m_{h^\pm} close to m_W , $e^+ e^- \rightarrow h^+ h^- \rightarrow \tau^+ \tau^- 2\nu_\tau$ could explain the 2.8σ deviation from lepton universality in W decays measured at LEP.
4. $B(h_2 \rightarrow a_1 a_1)$ and/or $B(h_2 \rightarrow Z a_1)$ are large. Thus, even if $e^+ e^- \rightarrow Z h_2$ has large σ (which is often the case since m_{h_2} is not large), would not have seen it since the $h_2 \rightarrow Z a_1$ decay was never looked for and an incomplete job was done on $h_2 \rightarrow a_1 a_1 \rightarrow 4\tau$.
5. For $\tan\beta = 1.7$ it is easy to find cases where $e^+ e^- \rightarrow Z h_1 \rightarrow Z b \bar{b}$ and $e^+ e^- \rightarrow Z h_2 \rightarrow Z b \bar{b}$ would yield a substantial contribution to the LEP $(0.1 - 0.2) \times SM$ excess near $m_{b\bar{b}} \sim 98$ GeV.
6. To observe or constrain the a_1 for larger (light- a finetuning preferred) $m_{a_1} \lesssim 2m_B$, will require Tevatron high luminosity data or LHC. **Still lots of models, even if not all, can be probed in this way.**
7. High Tevatron L would also better limit $B(t \rightarrow h^+ b)$ which at the moment is allowed up to the 40% level as these decays are included in the way CDF and D0 determine the $t\bar{t}$ cross section for the $h^+ \rightarrow W^+ a_1$.

Detecting the light h of the NMSSM

LHC

All standard LHC channels fail: *e.g.* $B(h \rightarrow \gamma\gamma)$ is much too small because of large $B(h \rightarrow aa)$.

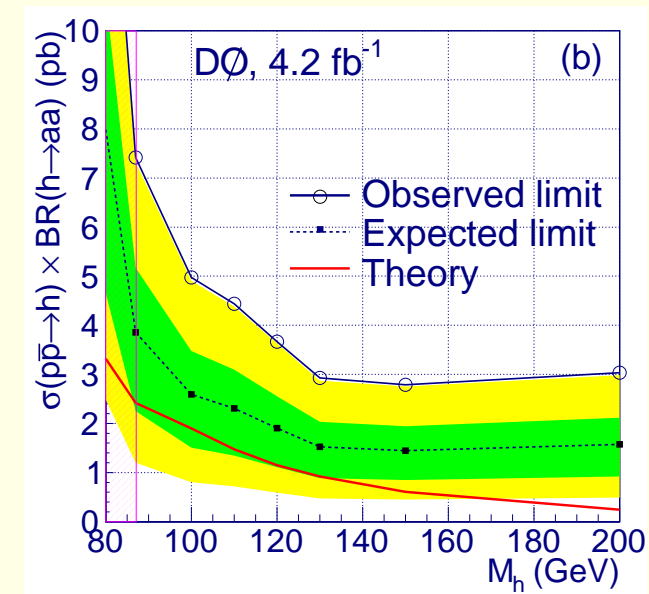
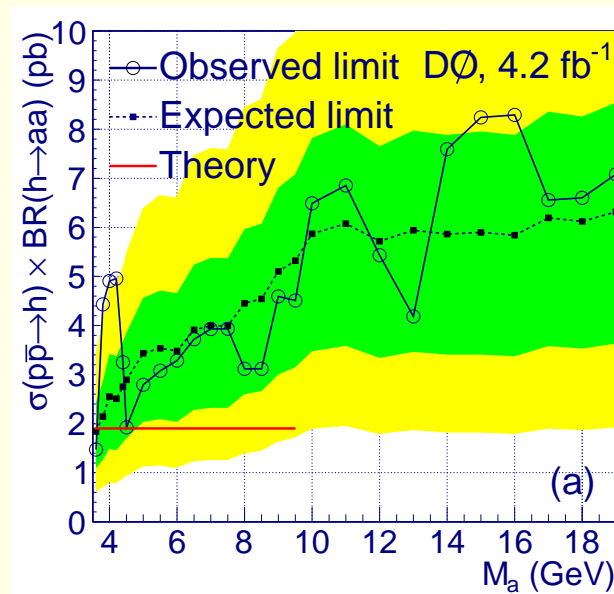
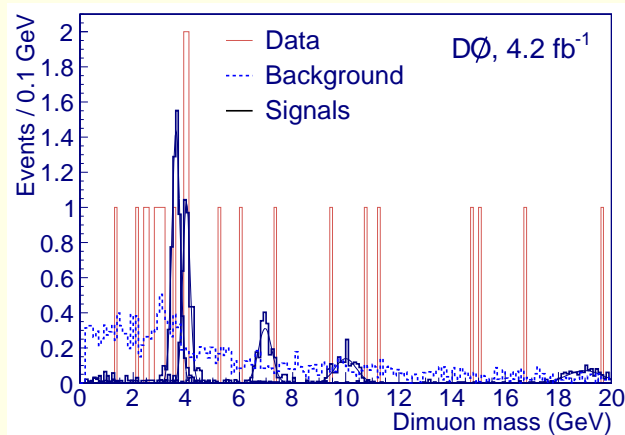
The possible new LHC channels include:

1. $gg \rightarrow h \rightarrow aa \rightarrow 4\tau$ **and** $2\tau + \mu^+\mu^-$

Always use μ tag for accepted events. $2\tau + 2\mu$ is main signal source after cuts.

There is an actual D0 analysis (A. Haas et. al.) of this mode using about $L \sim 4 \text{ fb}^{-1}$ of data. There are even small $\sim 1\sigma$ excesses for $m_a \sim 4$ and $10 - 11 \text{ GeV}$ consistent with predicted signal. About $L \sim 40 \text{ fb}^{-1}$ would

be needed for a 3σ signal.



From arXiv:0905.3381.

At the LHC? Studied by Wacker et al.

- $\sigma(gg \rightarrow h) \sim 50$ pb for $m_h \sim 100$ GeV.
- $B(h \rightarrow aa) \sim 0.8 - 0.9$.
- $B(a \rightarrow \mu^+\mu^-) \sim 0.0035 - 0.004$ and $B(a \rightarrow \tau^+\tau^-) \sim 0.95 - 0.98$
- Useful branching ratio product is $2 \times B(a \rightarrow \mu^+\mu^-)B(a \rightarrow \tau^+\tau^-) \sim .0075$.
- Cut efficiencies $\epsilon \sim 0.018$.

- Net useful cross section:

$$\sigma(gg \rightarrow h)B(h \rightarrow aa)[2B(a \rightarrow \mu^+\mu^-)B(a \rightarrow \tau^+\tau^-)]\epsilon \sim 4 - 7 \text{ fb}. \quad (2)$$

Backgrounds are small so perhaps 10 events in a single $\mu^+\mu^-$ bin would be convincing \Rightarrow need about $L = 2 \text{ fb}^{-1}$.

Note: If $m_a < 2m_\tau$, then $B(a \rightarrow \mu^+\mu^-) > 0.06$ and

$$\sigma(gg \rightarrow h)B(h \rightarrow aa)[B(a \rightarrow \mu^+\mu^-)]^2\epsilon > (153 \text{ fb}) \times \epsilon. \quad (3)$$

If $\epsilon > 0.02$ (seems likely) then $\Rightarrow \sigma_{eff} > 3 \text{ fb}$. This should be really background free and would close the $m_a < 2m_\tau$ "window of worry".

2. $WW \rightarrow h \rightarrow aa \rightarrow \tau^+\tau^- + \tau^+\tau^-$.

Key will be to tag relevant events using spectator quarks and require very little activity in the central region by keeping only events with 4 or 6 tracks.

Looks moderately promising but far from definitive results at this time (see, A. Belyaev *et al.*, arXiv:0805.3505 [hep-ph] and our work, JFG+Tait+Z. Han, below).

More shortly.

3. $t\bar{t}h \rightarrow t\bar{t}aa \rightarrow t\bar{t} + \tau^+\tau^- + \tau^+\tau^-.$

No study yet. Would isolated tracks/leptons from τ 's make this easier than $t\bar{t}h \rightarrow t\bar{t}b\bar{b}$?

4. $W, Z + h \rightarrow W, Z + aa \rightarrow W, Z + \tau^+\tau^- + \tau^+\tau^-.$

Leptons from W, Z and isolated tracks/leptons from τ 's would provide a clean signal. No study yet.

5. $\tilde{\chi}_2^0 \rightarrow h\tilde{\chi}_1^0$ with $h \rightarrow aa \rightarrow 4\tau.$

(Recall that the $\tilde{\chi}_2^0 \rightarrow h\tilde{\chi}_1^0$ channel provides a signal in the MSSM when $h \rightarrow b\bar{b}$ decays are dominant.)

6. **Last, but definitely not least: diffractive production $pp \rightarrow pph \rightarrow ppX.$**

The mass M_X can be reconstructed with roughly a 1 – 2 GeV resolution, potentially revealing a Higgs peak, independent of the decay of the Higgs.

The event is quiet so that the tracks from the τ 's appear in a relatively clean environment, allowing track counting and associated cuts.

Signal significances from JFG, Forshaw, Pilkington, Hodgkinson, Papaefstathiou: arXiv:0712.3510 are plotted in Fig. 18 for a variety of luminosity and triggering assumptions.

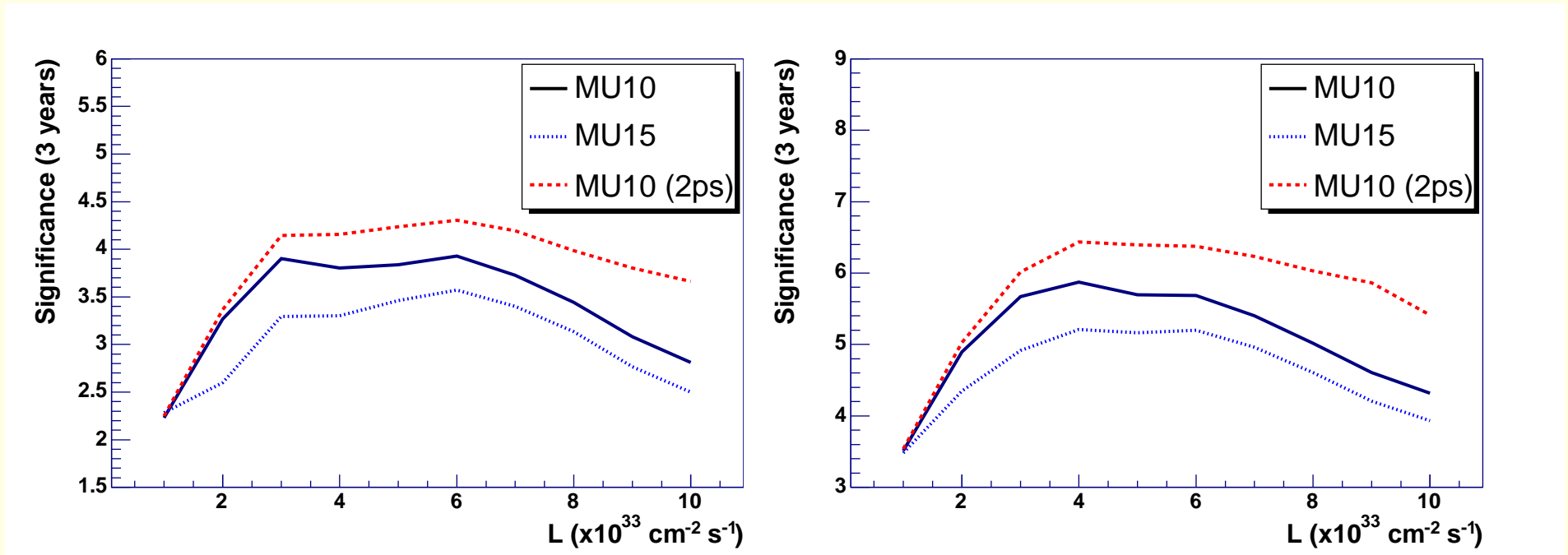


Figure 18: (a) The significance for three years of data acquisition at each luminosity. (b) Same as (a) but with twice the data. Different lines represent different μ trigger thresholds and different forward detector timing. Some experimentalists say more efficient triggering is possible, doubling the number of events at given luminosity.

CMS folk claim we can increase our rates by about a factor of 2 to 3 using additional triggering techniques.

The Collinearity Trick

- Since $m_a \ll m_h$, the a 's in $h \rightarrow aa$ are highly boosted.
 \Rightarrow the a decay products will travel along the direction of the source a .
 $\Rightarrow p_a \propto \sum$ visible 4-momentum of the charged tracks in its decay.
Labeling the two a 's with indices 1 and 2 we have

$$p_i^{vis} = f_i p_{a,i} \quad (4)$$

where $1 - f_i$ is the fraction of the a momentum carried away by neutrals.

- $pp \rightarrow pph$ case

The accuracy of this has now been tested in the $pp \rightarrow pph$ case, and gives an error for m_h of order 5 GeV, but this is less accurate than m_h determination from the tagged protons and so is not used.

However, we are able to make *four* m_a determinations per event.

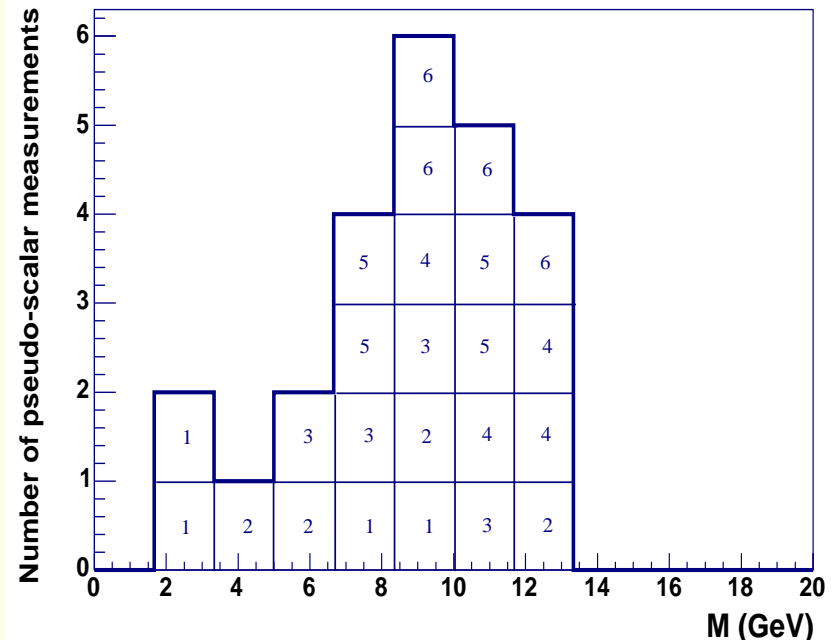
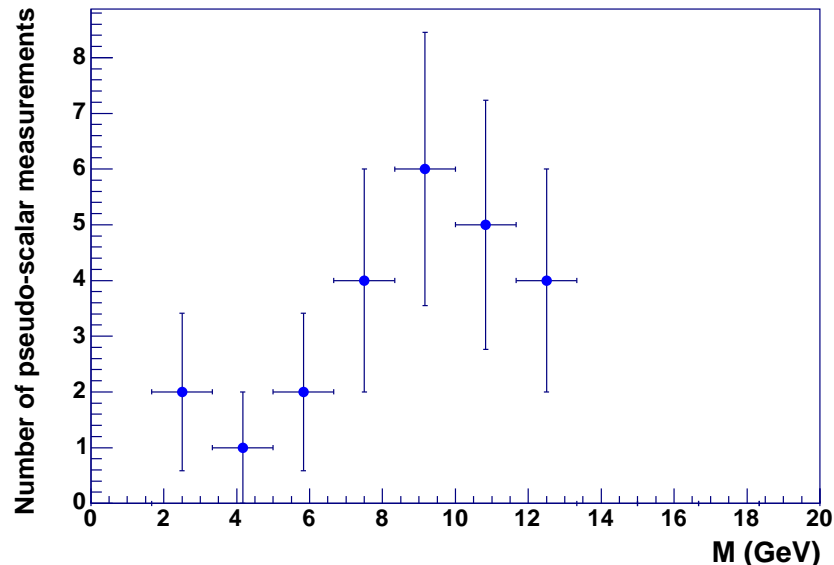


Figure 19: (a) A typical a mass measurement. (b) The same content as (a) but with the breakdown showing the 4 Higgs mass measurements for each of the 6 events, labeled 1 – 6 in the histogram.

Figure 19 shows the distribution of masses obtained for 180 fb^{-1} of data collected at $3 \times 10^{33} \text{ cm}^{-2}\text{s}^{-1}$, corresponding to about 6 Higgs events and therefore 24 m_a entries.

By considering many pseudo-data sets, we conclude that a typical experiment would yield $m_a = 9.3 \pm 2.3 \text{ GeV}$, which is in re-assuringly good agreement with the input value of 9.7 GeV.

- $WW \rightarrow h$

For $m_h = 100$ GeV and SM-like $WW h$ coupling, $\sigma(WW \rightarrow h) \sim 7$ pb, implying 7×10^5 events before cuts for $L = 100$ fb $^{-1}$.

In this case, we do not know the longitudinal momentum of the h , but we should have a good measurement of its transverse momentum from the tagging jets and other recoil jets.

This gives two equations in the two unknown $f_{1,2}$ and allows us to solve and construct mass peaks.

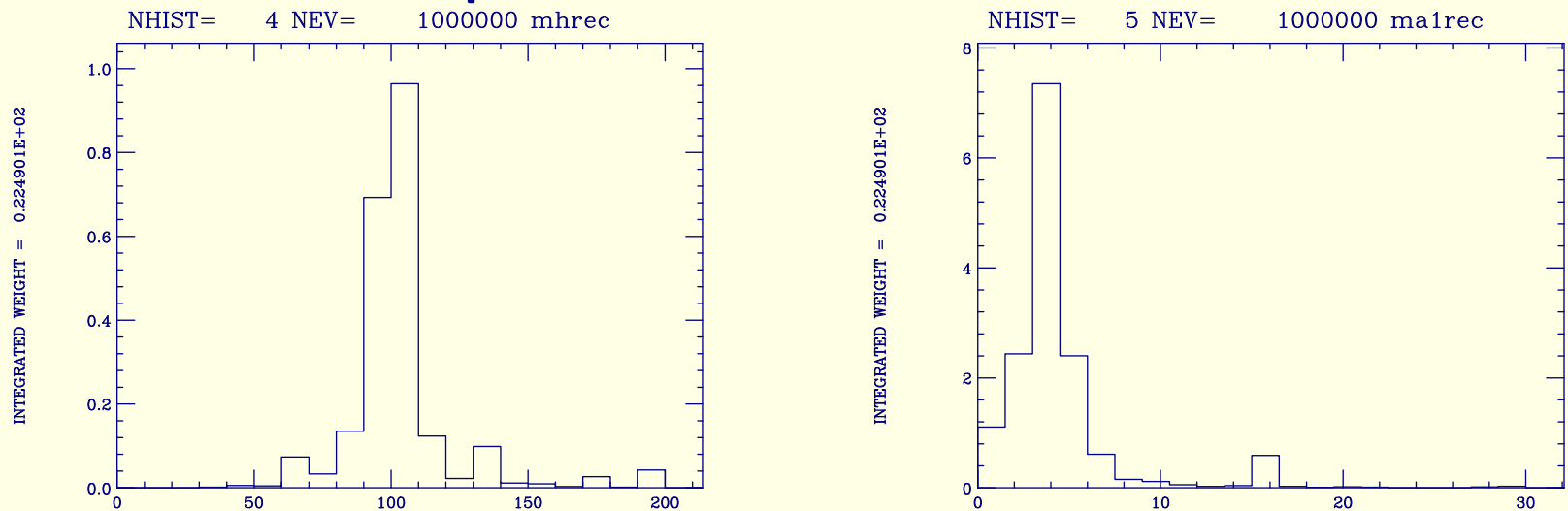


Figure 20: (a) A typical h mass distribution. (b) A typical a mass distribution. No cuts imposed; signal only.

Other related scenarios

- A string of Higgs, as possibly hinted at by the CDF multi-muon events.

The SM-like Higgs could then decay into a string of Higgs bosons: e.g.
 $h \rightarrow h_1 h_1 \rightarrow (h_2 h_2)(h_2 h_2) \rightarrow ((h_3 h_3)(h_3 h_3))((h_3 h_3)(h_3 h_3)) \rightarrow \dots$
(Any of the h_i 's could be a 's and then $a_i \rightarrow a_j h_k$ would follow.)

(Ellwanger et al have an NMSSM model that gives CDF multi-muon, but implications for unusual h decays are unclear.)

- Many singlets, as generically possible in string models, could mix with the doublet Higgs and create a series of Higgs eigenstates (with mass weight in the < 100 GeV region for good PEW).

It can be arranged that these eigenstates decay in complex ways that would have escaped LEP limits.

In fact, one can get really low "effective" Higgs mass from PEW point of view while fitting under LEP constraint curve.

This is the "worst case" scenario envisioned long ago in JFG, Espinosa: hep-ph/9807275.

- Low $\tan\beta$ NMSSM scenarios in which the first two CP-even Higgs bosons both have mass in the $\lesssim 100$ GeV region and decay so as to escape LEP (and Tevatron) limits. See earlier section.
- Drop dark matter requirement: \Rightarrow huge plethora of possibilities in SUSY. Includes "hidden valley" decays, R -parity violating decays,

ILC

At the ILC, there is no problem since $e^+e^- \rightarrow ZX$ will reveal a $M_X \sim m_h \sim 90 - 100$ GeV peak no matter how the h decays.

If there are many Higgs, then the excesses in various bins of M_X will be apparent even if there is a broad sort of spectrum and X has a mixture of decays.

But the ILC is decades away.

Conclusions

In case you hadn't noticed, theorists have been going a bit crazy waiting for the Higgs.



"Unfortunately", a lot of the theories developed make sense, but I remain enamored of the NMSSM scenarios and hope for eventual verification that nature has chosen "wisely".

Meanwhile, all I can do is watch and wait (but perhaps not from quite so close a viewpoint).

