

# Higgs-Radion mixing and the LHC Higgs-like excesses

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# Higgs-like LHC Excesses

Is what we are seeing a Higgs-like chameleon?



# Higgs-like LHC Excesses

Or is it **THE** Higgs?



- Given that the mass(es) is(are) of order 125 GeV, the MSSM or the much more attractive NMSSM extension thereof is a natural candidate theory.

After all, SUSY solves the hierarchy problem, predicts gauge coupling unification at the GUT scale and so forth.

However:

- A SM-like Higgs with mass as large as 125 GeV is a bit of a stretch. Even in the NMSSM 125 GeV is “on the edge” for semi-universal (NUHM) GUT boundary conditions (and not possible for full CNMSSM b.c.).
  - This is aggravated if the signal is  $> SM$ .
  - And, even more problematically, there may be more than one ‘excess’ in the data (cf. CMS data).
- The only other really attractive alternate solution to the hierarchy problem that provides a self-contained ultraviolet complete framework is to allow extra dimensions.

One particular implementation is the Randall Sundrum model in which there is a warped 5th dimension.

# The Randall Sundrum Model

- The background RS metric that solves Einstein's equations takes the form[1]

$$ds^2 = e^{-2m_0 b_0 |y|} \eta_{\mu\nu} dx^\mu dx^\nu - b_0^2 dy^2 \quad (1)$$

where  $y$  is the coordinate for the 5th dimension with  $|y| \leq 1/2$ .

- The graviton and radion fields,  $h_{\mu\nu}(x, y)$  and  $\phi_0(x)$ , are the quantum fluctuations relative to the background metric  $\eta_{\mu\nu}$  and  $b_0$ , respectively.
- In the simplest case, only gravity propagates in the bulk while the SM is located on the infrared (or TeV) brane at  $y = 1/2$  and

$$\mathcal{L}_{\text{int}} = -\frac{1}{\widehat{\Lambda}_W} \sum_{n \neq 0} h_{\mu\nu}^n T^{\mu\nu} - \frac{\phi_0}{\Lambda_\phi} T^\mu_\mu \quad (2)$$

where  $h_{\mu\nu}^n(x)$  are the Kaluza-Klein (KK) modes (with mass  $m_n$ ) of the graviton field  $h_{\mu\nu}(x, y)$ .

- $\Lambda_\phi$  is the vacuum expectation value of the radion field and is related to  $\hat{\Lambda}_W$  and  $m_{Pl}$  as follows.

$$\hat{\Lambda}_W \simeq \sqrt{2} m_{Pl} \Omega_0, \quad \text{where} \quad \Omega_0 = e^{-\frac{1}{2} m_0 b_0}, \quad \Lambda_\phi = \sqrt{3} \hat{\Lambda}_W. \quad (3)$$

- If matter propagates in the bulk then the interactions of gravitons and radion with matter are controlled by the overlap of appropriate extra-dimensional profiles and corrections to (2) appear.
- Besides the radion, the model contains a conventional Higgs boson,  $h_0$ .
- The RS model provides a simple solution to the hierarchy problem if the Higgs is placed on the TeV brane at  $y = 1/2$  by virtue of the fact that the 4D electro-weak scale  $v_0$  is given in terms of the  $\mathcal{O}(m_{Pl})$  5D Higgs vev,  $\hat{v}$ ,

by:

$$v_0 = \Omega_0 \hat{v} = e^{-\frac{1}{2} m_0 b_0} \hat{v} \sim 1 \text{ TeV} \quad \text{for} \quad \frac{1}{2} m_0 b_0 \sim 35. \quad (4)$$

To solve the hierarchy problem,  $\Lambda_\phi = \sqrt{6} m_{Pl} \Omega_0 \lesssim 1 - 10 \text{ TeV}$  needed.

- $m_0/m_{Pl}$  is a particularly crucial parameter that characterizes the 5-dimensional curvature.

$m_0/m_{Pl} \gtrsim 0.5$  is favored for fitting the LHC Higgs excesses and by bounds on FCNC and PEW constraints.

Views on  $m_0/m_{Pl}$  are changing:

- Original:  $R_5/M_5^2 < 1$  ( $M_5$  being the 5D Planck scale and  $R_5 = 20m_0^2$  the size of the 5D curvature) is needed to suppress higher curvature terms in the 5D action:  $\Rightarrow m_0/m_{Pl} \lesssim 0.15$ .
- New: [9] argues that  $R_5/\Lambda^2$  ( $\Lambda$  = energy scale at which the 5D gravity theory becomes strongly coupled, with NDA estimate of  $\Lambda \sim 2\sqrt{3}\pi M_5$ ), is the appropriate measure,  
 $\Rightarrow m_0/m_{Pl} < \sqrt{3\pi^3/(5\sqrt{5})} \sim 3$  acceptable.

- Note: the mass of 1st KK graviton excitation ( $G^1$ ) is related to  $m_0/m_{Pl}$  and  $\Lambda_\phi$  by

$$m_1^{\text{KK}} = \frac{(m_0/m_{Pl})x_1^{\text{KK}}}{\sqrt{6}}\Lambda_\phi, \quad (5)$$

where  $x_1^{\text{KK}} \sim 3.83$ .

$\Rightarrow$  large  $m_0/m_{Pl}$  if the lower bound on  $m_1^{\text{KK}}$  is large and  $\Lambda_\phi \sim 1$  TeV.

- In the simplest RS scenario, the SM fermions and gauge bosons are confined to the brane.

Now regarded as highly problematical:

- Higher-dimensional operators in the 5D effective field theory are suppressed only by  $\text{TeV}^{-1}$ ,  $\Rightarrow$  FCNC processes and PEW observable corrections are predicted to be much too large.  
 $\Rightarrow$  no explanation of the flavor hierarchies.

- Must move fermions and gauge bosons (but not Higgs) off the brane  
[\[2\]](#)[\[3\]](#)[\[4\]](#)[\[5\]](#)[\[6\]](#)[\[7\]](#)[\[8\]](#)[\[9\]](#)



The SM particles = zero-modes of the 5D fields and the profile of a SM fermion in the extra dimension can be adjusted using a mass parameter.

Two possibilities:

1. If 1st and 2nd generation fermion profiles peak near the Planck brane then FCNC operators and PEW corrections will be suppressed by scales  $\gg \text{TeV}$ .

Even with this arrangement it is estimated that the  $g^1$ ,  $W^1$  and  $Z^1$  masses must be larger than about 3 TeV (see the summary in [9]).

2. Use fairly flat profiles for the 1st and 2nd generation fermions in the 5th dimension.

$\Rightarrow$  the coupling of light quarks to  $g^1$ ,  $W^1$ ,  $Z^1$ , proportional to the integral of the square of the fermion profile multiplied by the gauge boson profile, will be very small, implying small FCNC.

PEW constraints would still be very problematical unless a special 5D GIM mechanism is employed [11] and/or there is a custodial symmetry in the bulk and there is a discrete  $L \leftrightarrow R$  symmetry to suppress  $Z \rightarrow b\bar{b}$  corrections.

In either case, the interactions of Eq. (2) are greatly modified when gauge bosons and fermions are allowed to propagate in the bulk.

- If the gauge bosons and fermions do not propagate in the bulk, then the strongest limits on  $\Lambda_\phi$  come, via Eq. (5), from the lower bound placed by the LHC on the first graviton KK excitation (see, for example, [13] and [14] for the ATLAS and CMS limits).
- However, when the fermions propagate in the bulk, the couplings of light fermion pairs to  $G^1$  are greatly reduced and these limits do not apply.
- When gauge bosons propagate in the bulk, a potentially important experimental limit on the model comes from lower bounds on the 1st excitation of the gluon,  $g^1$ .
- **First approach:** In the model of [12], in which light fermion profiles peak near the Planck brane, there is a universal component to the light quark coupling  $q\bar{q}g^1$  that is roughly equal to the SM coupling  $g$  times a factor of  $\zeta^{-1}$ , where  $\zeta \sim \sqrt{\frac{1}{2}m_0 b_0} \sim 5 - 6$ .

The suppression is due to the fact that the light quarks are localized near the Planck brane whereas the KK gluon is localized near the TeV brane.

Even with such suppression, the LHC  $g^1$  production rate due to  $u\bar{u}$  and  $d\bar{d}$  collisions is large.

Further, the  $t_R\bar{t}_R g^1$  coupling is large since the  $t_R$  profile peaks near the TeV brane – the prediction of [12] is  $g_{t_R\bar{t}_R g^1} \sim \zeta g$ .

$\Rightarrow g^1 \rightarrow t\bar{t}$  decays dominant.

$\Rightarrow$  lower bound of  $m_1^g \gtrsim 1.5$  TeV [15] using an update of the analysis of [12]. ([16] gives a weaker bound of  $m_1^g > 0.84$  TeV.) .

- In terms of  $\Lambda_\phi$ , we have the following relations:

$$\frac{m_0}{m_{Pl}} = \frac{\sqrt{6} m_1^g}{x_1^g \Lambda_\phi} \simeq \frac{m_1^g}{\Lambda_\phi}, \quad \text{and} \quad \frac{1}{2} m_0 b_0 = -\log \left( \frac{\Lambda_\phi}{\sqrt{6} m_{Pl}} \right) \quad (6)$$

where  $x_1^g = 2.45$  is the 1st zero of an appropriate Bessel function.

If the model really solves the hierarchy problem then  $\hat{\Lambda}_W$  in Eq. (2) cannot be much larger than 1 TeV. Let us take  $\hat{\Lambda}_W \lesssim 5.5$  TeV as an extreme possibility, which implies  $\Lambda_\phi = \sqrt{3}\hat{\Lambda}_W \lesssim 10$  TeV.

If we adopt the CMS limit of  $m_1^g > 1.5$  TeV then (6) implies a lower limit on the 5-dimensional curvature of  $m_0/m_{Pl} \gtrsim 0.15$ .

Thus, a significant lower bound on  $m_1^g$  implies that only relatively large values for  $m_0/m_{Pl}$  are allowed.  $\Rightarrow$  probably ok given latest ideas.

- **Alternate approach:** Take light fermion profiles to be flat in 5th dimension.  
 $\Rightarrow q\bar{q}g^1, \dots$  couplings are small (fix precision data as noted earlier)  
 $\Rightarrow$  no direct experimental bound on  $m_1^g$ .

$\Lambda_\phi$ ?

if  $\Lambda_\phi = 1$  TeV, for  $m_0/m_{Pl} = 0.01, 0.1$  Eq. (6) implies  $m_1^g = 10, 100$  GeV.

Could the  $g^1$  really have escaped discovery for such low masses?

If there is no firm bound on  $m_1^g \Rightarrow$  discuss the phenomenology for fixed  $\Lambda_\phi$ . We will consider  $\Lambda_\phi = 1$  TeV and 1.5 TeV.

## Higgs-Radion Mixing

- Since the radion and higgs fields have the same quantum numbers, they can mix. [17]

$$S_\xi = \xi \int d^4x \sqrt{g_{\text{vis}}} R(g_{\text{vis}}) \widehat{H}^\dagger \widehat{H}, \quad (7)$$

The physical mass eigenstates,  $h$  and  $\phi$ , are obtained by diagonalizing and canonically normalizing the kinetic energy terms.

The diagonalization procedures and results for the  $h$  and  $\phi$  using our notation can be found in [10] (see also [17][18]).

In the end, one finds

$$h_0 = dh + c\phi \quad - \quad \phi_0 = a\phi + bh. \quad (8)$$

where  $a, b, c, d$  are functions of the eigenstate masses,  $m_h$  and  $m_\phi$ ,  $\xi$  and  $\gamma \equiv v_0/\Lambda_\phi$  ( $v_0 = 246$  GeV). For  $\xi = 0$ ,  $d = 1, a = -1, c = b = 0$ . Radion couplings to SM  $\propto \gamma$ .

Consistency of the diagonalization  $\Rightarrow$  strong restrictions on the  $\xi$  range as a function of the final eigenstate masses  $m_h$  and  $m_\phi$

- For the full Feynman rules for the  $h$  and  $\phi$  when fermions and gauge bosons propagate in the bulk see [19].

Of particular note are the anomaly terms associated with the  $\phi_0$  interactions before mixing. Defining

$$g_h = (d + \gamma b) \quad g_\phi = (c + \gamma a) \quad g_h^r = \gamma b \quad g_\phi^r = \gamma a. \quad (9)$$

$$c_g^{h,\phi} = -\frac{\alpha_s}{4\pi v} \left[ g_{h,\phi} \sum_i F_{1/2}(\tau_i) - 2(b_3 + \frac{2\pi}{\alpha_s \frac{1}{2} m_0 b_0}) g_{h,\phi}^r \right]$$

$$c_\gamma^{h,\phi} = -\frac{\alpha}{2\pi v} \left[ g_{h,\phi} \sum_i e_i^2 N_c^i F_i(\tau_i) - (b_2 + b_Y + \frac{2\pi}{\alpha \frac{1}{2} m_0 b_0}) g_{h,\phi}^r \right] \quad (10)$$

New relative to old on-the-brane results are the  $2\pi \dots$  corrections to the  $g^r$  terms. They can be significant in the  $c_\gamma$  case.

- There are also modifications to the  $WW$  and  $ZZ$  couplings of the  $h$  and  $\phi$  relative to old on-the-brane results.

These bring in  $W_{\mu\nu}^\dagger W^{\mu\nu}$  terms in  $\mathcal{L}$  in addition to the usual  $W_\mu^\dagger W^\mu$  terms.

## LHC Excesses

- In the context of the higgs-radion model, positive signals can only arise for two masses.
- If more than two excesses were to ultimately emerge, then a more complicated Higgs sector will be required than the single  $h_0$  case we study here.

Certainly, one can consider including extra Higgs singlets or doublets.

For the moment, we presume that there are at most two excesses. In this case, it is sufficient to pursue the single Higgs plus radion model.

We will consider a few cases. Errors quoted for the excesses are  $\pm 1\sigma$ .

### 1. ATLAS:

125 GeV:  $\gamma\gamma$  excess of  $2_{-0.8}^{+0.8} \times \text{SM}$

125 GeV:  $4\ell$  excess of  $1.5_{-1}^{+1.5} \times \text{SM}$

### 2. CMSA:

124 GeV:  $\gamma\gamma$  excess of  $1.7_{-0.7}^{+0.8} \times \text{SM}$

124 GeV:  $4\ell$  excess of  $0.5_{-0.7}^{+1.1} \times \text{SM}$

120 GeV:  $4\ell$  excess of  $\sim 2_{-1}^{+1.5} \times \text{SM}$  but  $\gamma\gamma$  rate  $< 0.5 \times \text{SM}$ .

### 3. CMSB:

124 GeV: as above

137 GeV:  $\gamma\gamma$  excess of  $1.5_{-0.8}^{+0.8} \times \text{SM}$

137 GeV:  $4\ell < 0.2 \times \text{SM}$

### Notes:

- For plots use 125 GeV always: no change if  $125 \text{ GeV} \rightarrow 124 \text{ GeV}$ .
- As discussed, consider two different kinds of models:
  1. lower bound on  $m_1^g$  of 1.5 TeV
  2. fixed  $\Lambda_\phi$



## Lower bound of $m_1^g = 1.5$ TeV

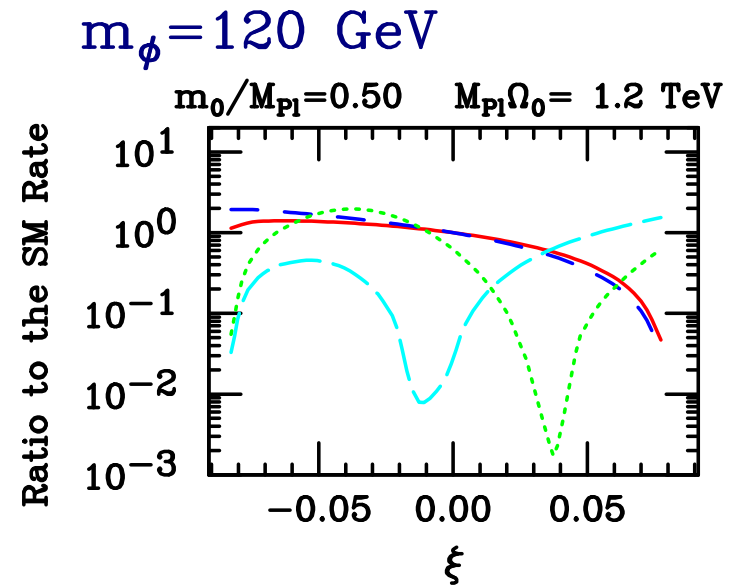
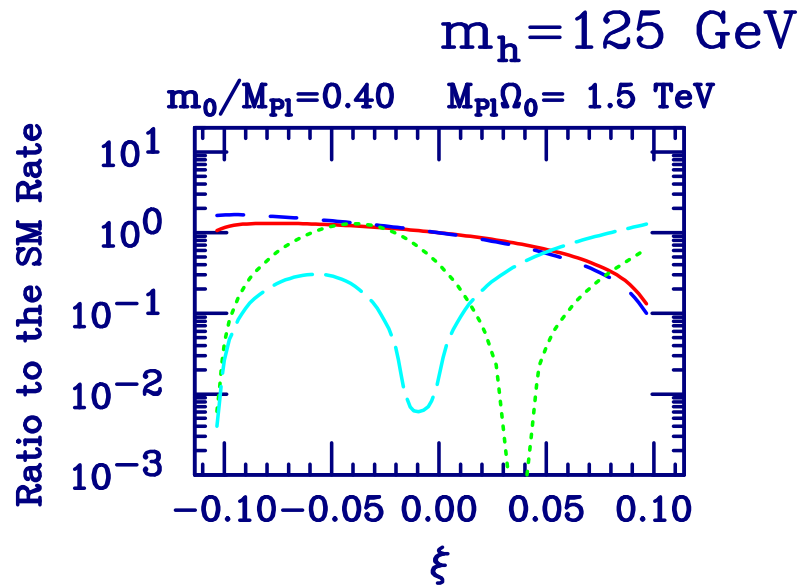
- Recall that  $\Lambda_\phi$  will be correlated with  $m_0/m_{Pl}$ .

$$\frac{m_0}{m_{Pl}} \simeq \frac{m_1^g}{\Lambda_\phi} \quad (11)$$

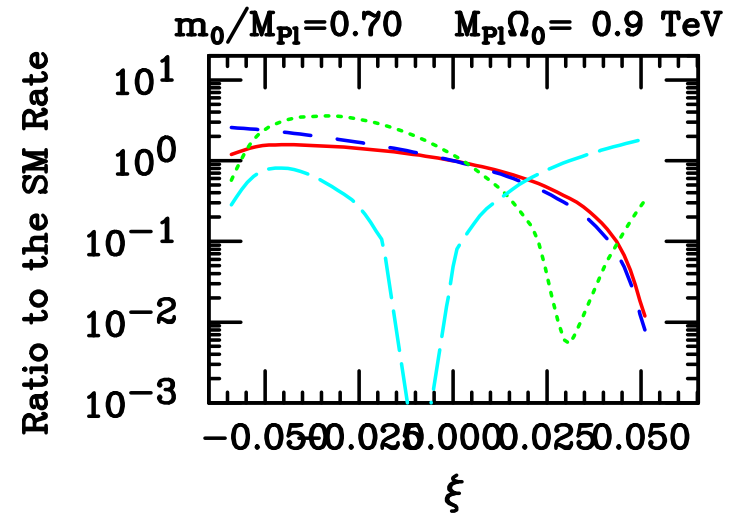
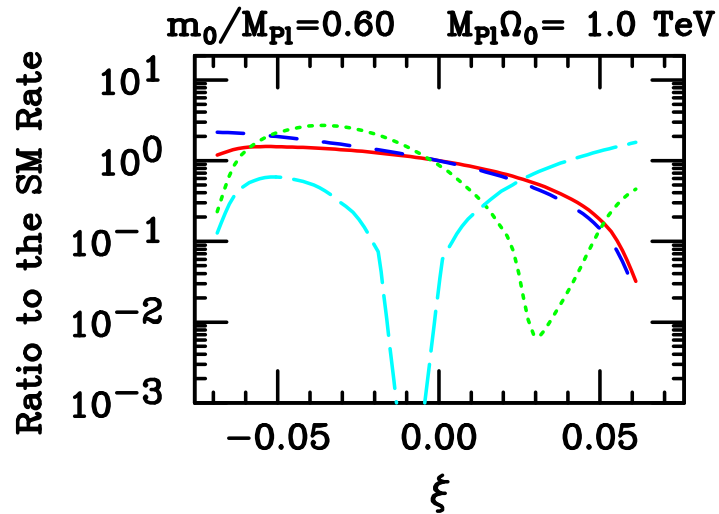
$\Rightarrow$

For small  $m_0/m_{Pl}$ ,  $\Lambda_\phi$  is too large, so only solve hierarchy if  $m_0/m_{Pl}$  is  $\gtrsim 0.2$ .

- Only have time for a limited selection of situations.



$h \rightarrow \gamma\gamma$ : solid red;  $h \rightarrow ZZ$ : blue dashes;  $\phi \rightarrow \gamma\gamma$ : green dots;  $\phi \rightarrow ZZ$ : cyan long dashes



**Figure 1:** We plot  $\gamma\gamma$  and  $ZZ$  relative to SM vs  $\xi$ .

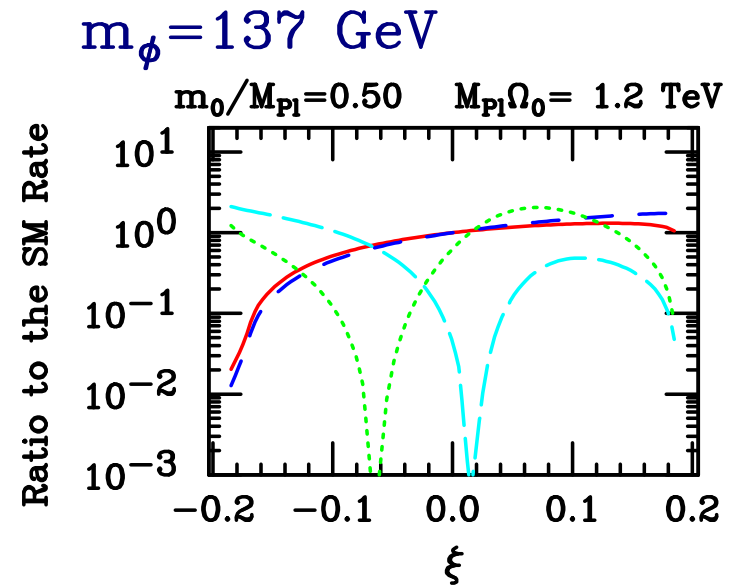
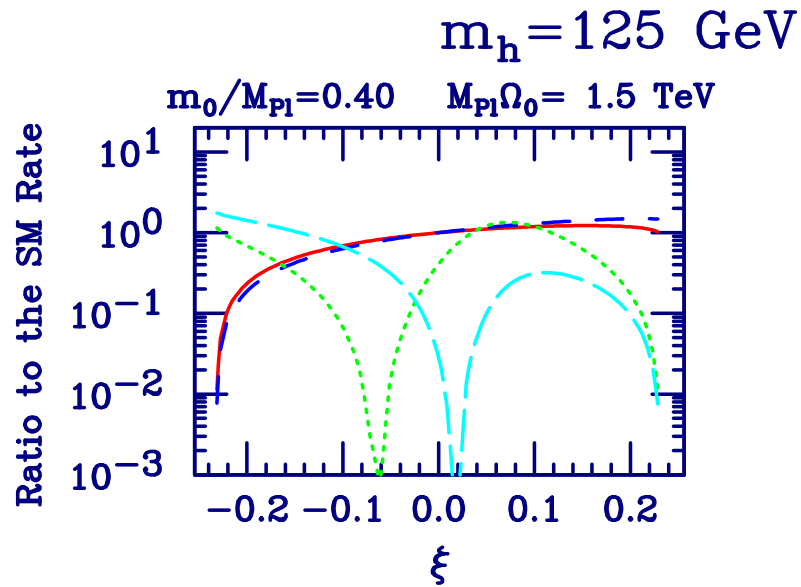
### Only 125 GeV excess

- If want no excesses at  $\sim 120$  GeV, but  $\gamma\gamma$  excess at 125 GeV of order  $\gtrsim 1.5 \times \text{SM}$ , then  $m_0/m_{Pl} = 0.4$  and  $\xi \sim -0.09$  are good choices.  
 $4\ell$  signal at 125 GeV is  $> \gamma\gamma$  but still within error.
- For the reversed assignments of  $m_h = 120$  GeV and  $m_\phi = 125$  GeV, no decent description of the ATLAS 125 GeV excesses with signals at 120 GeV being sufficiently suppressed.

### Excesses at 125 GeV and 120 GeV

- Higgs-radion scenario fails.

In the regions of  $\xi$  for which appropriate signals are present at 125 GeV from the  $h$ , the  $4\ell$  and  $\gamma\gamma$  rates at 120 GeV are either both suppressed or it is the  $\gamma\gamma$  rate that is enhanced more than the  $4\ell$  rate at 120 GeV. This phenomenon persists at higher  $m_0/m_{Pl}$  values.



$h \rightarrow \gamma\gamma$ : solid red;  $h \rightarrow ZZ$ : blue dashes;  $\phi \rightarrow \gamma\gamma$ : green dots;  $\phi \rightarrow ZZ$ : cyan long dashes

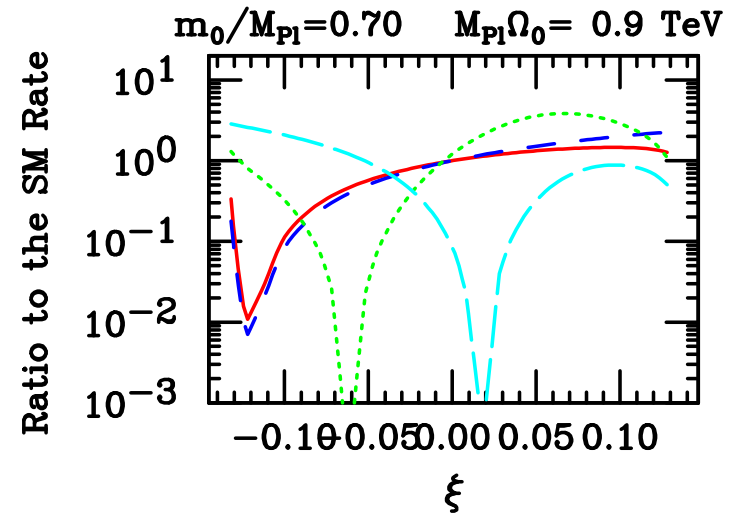
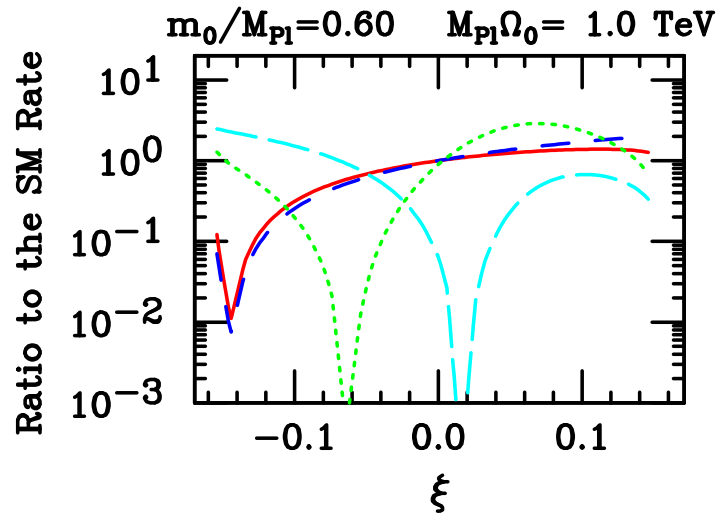


Figure 2:  $\gamma\gamma$  and  $ZZ$  relative to SM vs  $\xi$ .

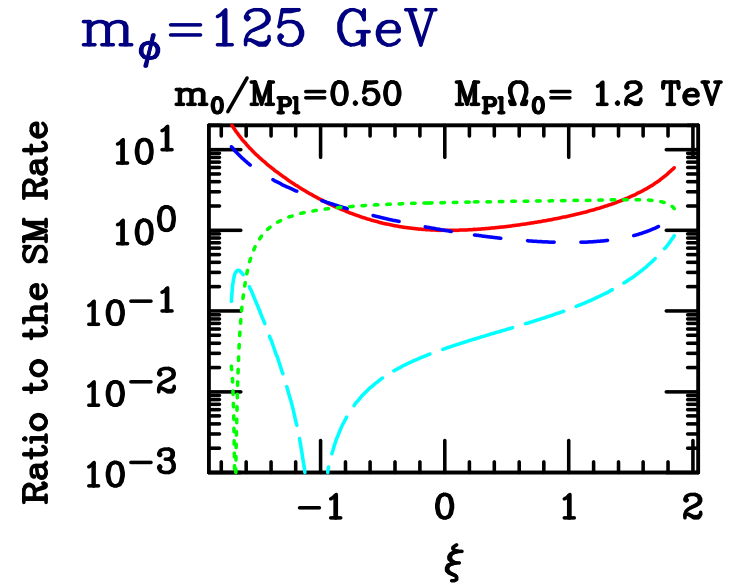
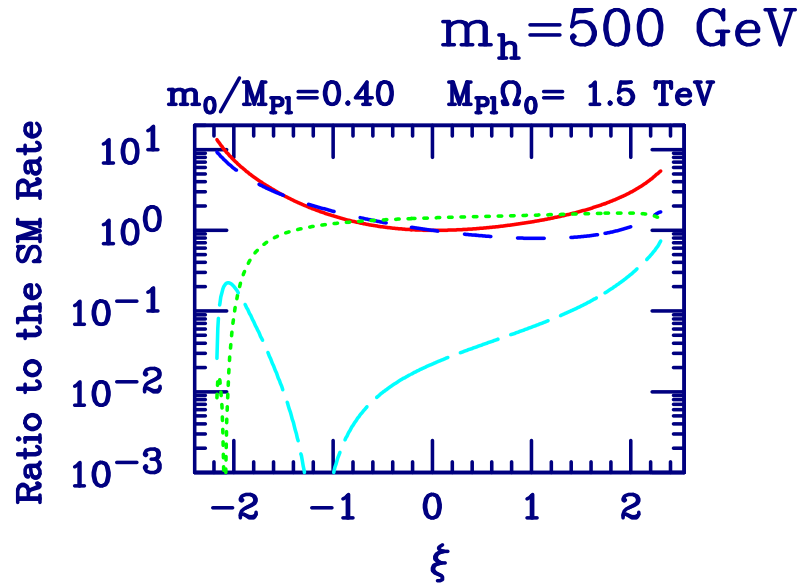
## Signals at 125 GeV and 137 GeV

- In Fig. 2 (previous page):  $m_0/m_{Pl} = 0.5$  and  $\xi = 0 \Rightarrow$

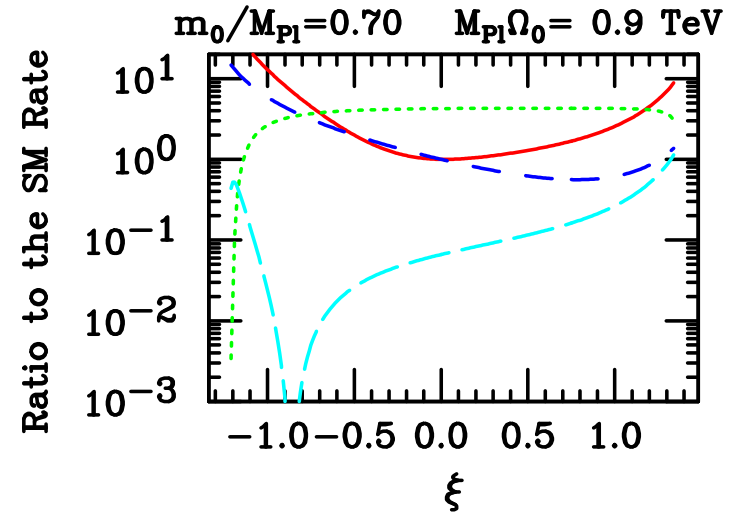
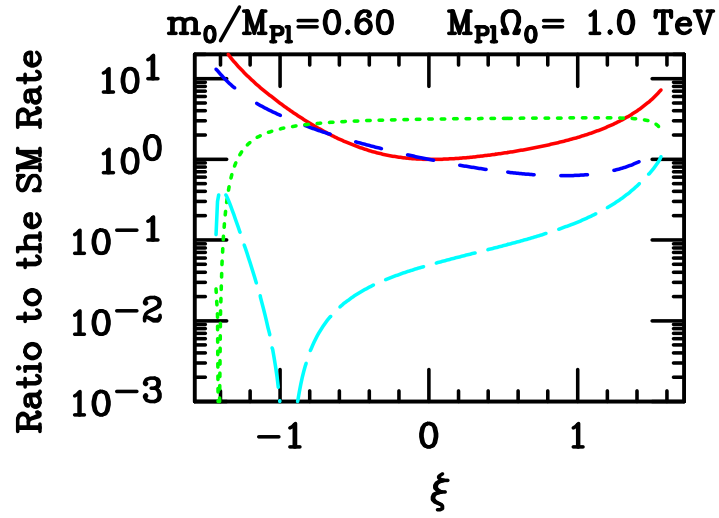
125 GeV:  $\gamma\gamma \sim 1 \times \text{SM}$  and  $4\ell \sim 1 \times \text{SM}$

137 GeV:  $\gamma\gamma \sim 1 \times \text{SM}$  and  $4\ell$  very small

These rates are consistent within  $1\sigma$  with the CMS observations.



$h \rightarrow \gamma\gamma$ : solid red;  $h \rightarrow ZZ$ : blue dashes;  $\phi \rightarrow \gamma\gamma$ : green dots;  $\phi \rightarrow ZZ$ : cyan long dashes



**Figure 3:** We plot  $\gamma\gamma$  and  $ZZ$  relative to SM vs  $\xi$ .

## Signals at 125 GeV and 500 GeV

- There are two choices,  $m_h = 125$  GeV with  $m_\phi = 500$  GeV or reverse. Let us discuss the reverse. Fig. 3.
- $\xi \sim \xi^{\max} \Rightarrow \gamma\gamma \sim 2\times\text{SM}$  and  $4\ell \sim 1\times\text{SM}$  at  $m_\phi = 125$  GeV for  $m_0/m_{Pl} = 0.4$  and  $0.5$ .
- $h \rightarrow 4\ell$  signal at 500 GeV  $\sim\text{SM}$ .
- $\gamma\gamma$  signal at 500 GeV enhanced, but SM rate small, and thus almost surely unobservable.
- $4\ell$  rates at 500 GeV? ATLAS and CMS disagree.

Probably, the heavy  $h$  in this scenario would have to be placed beyond LHC reach (for which the  $m_\phi = 125$  GeV signals shown would be unchanged).

## Fixed $\Lambda_\phi$

- If fermionic profiles are quite flat, couplings of light quarks to the gauge excitations are very small.  $\Rightarrow$  no bounds on  $m_1^g$  or  $\Lambda_\phi$ .

We choose to examine the phenomenology for (low) values of  $\Lambda_\phi = 1$  TeV and  $\Lambda_\phi = 1.5$  TeV.

- When fermions and, in particular, gauge bosons propagate in the bulk the phenomenology does not depend on  $\Lambda_\phi$  alone — at fixed  $\Lambda_\phi$  there is strong dependence on  $m_0/m_{Pl}$  when  $m_0/m_{Pl}$  is small.
- Only for large  $m_0/m_{Pl} \gtrsim 0.5$  is the phenomenology determined almost entirely by  $\Lambda_\phi$ .

But, not the same as when all fields are on the TeV brane.

- Once again, we step through the various possible mass locations for the



Higgs and radion that are motivated by the LHC excesses in the  $\gamma\gamma$  and/or  $4\ell$  channels.

Single resonance at 125 GeV

- The choice of  $\Lambda_\phi = 1$  TeV with  $m_\phi = 125$  GeV and  $m_h = 120$  GeV gives a reasonable description of the ATLAS excesses at 125 GeV with no visible signals at 120 GeV in either the  $\gamma\gamma$  or  $4\ell$  channels. (figure not shown)

A good choice of parameters is  $m_0/m_{Pl} = 1$  and  $\xi = -0.015$ .

- In contrast, for  $\Lambda_\phi = 1.5$  TeV the 125 GeV predicted  $\gamma\gamma$  and  $4\ell$  excesses are below  $1 \times \text{SM}$  and thus would not provide a good description of the ATLAS excesses.
- For the reversed assignments of  $m_h = 125$  GeV and  $m_\phi = 120$  GeV, any choice of parameters that gives a good description of the 125 GeV signals yields a highly observable 120 GeV signal, not appropriate for ATLAS.

## Signals at 125 GeV and 120 GeV

- There is a marginal solution for  $m_h = 125$  GeV,  $m_\phi = 120$  GeV,  $\xi = \xi^{\max}$  and  $m_0/m_{Pl} = 1.1$ .
- With the reversed assignments of  $m_h = 120$  GeV and  $m_\phi = 125$  GeV a satisfactory description of the two CMS excesses is not possible.

## Signals at 125 GeV and 137 GeV

- CMS data want:

125 GeV:  $\gamma\gamma \sim 1.7 \times \text{SM}$  and  $4\ell < 1.6 \times \text{SM}$ .

137 GeV:  $\gamma\gamma \sim 1.5 \times \text{SM}$  and  $4\ell \sim \text{small}$ .

- For  $\Lambda_\phi = 1$  TeV,  $m_h = 125$  GeV,  $m_\phi = 137$  GeV get rough description at  $m_0/m_{Pl} = 0.6$  and  $\xi = -0.05$

- For  $\Lambda_\phi = 1.5$  TeV, can do better.

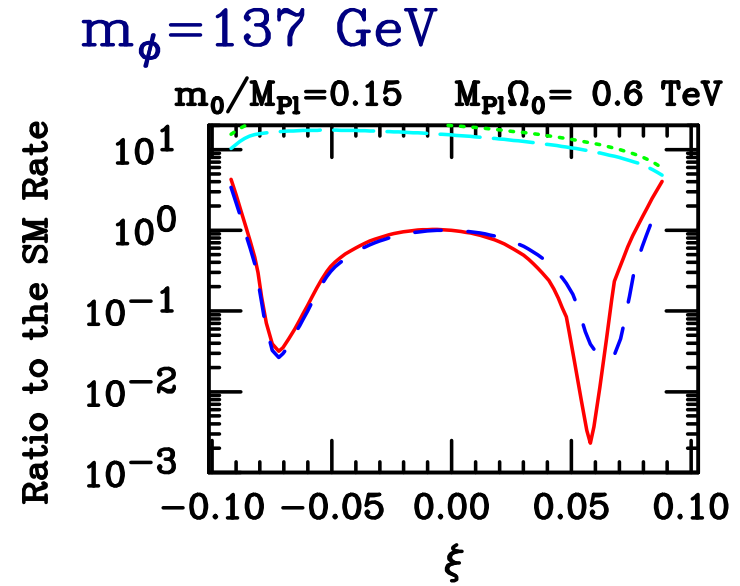
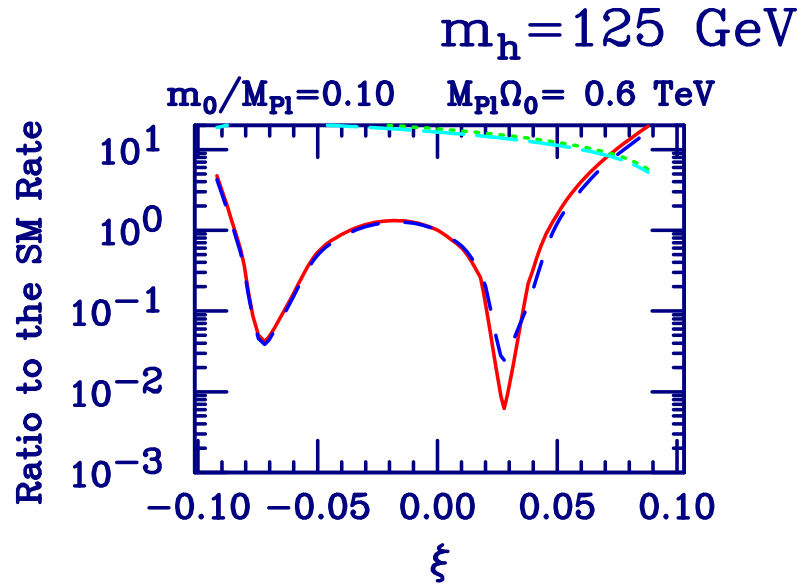
Fig. 4 (next page) shows results for  $m_h = 125$  GeV and  $m_\phi = 137$  GeV.

For  $m_0/m_{Pl} = 0.25$  and  $\xi \sim -0.1 \Rightarrow$

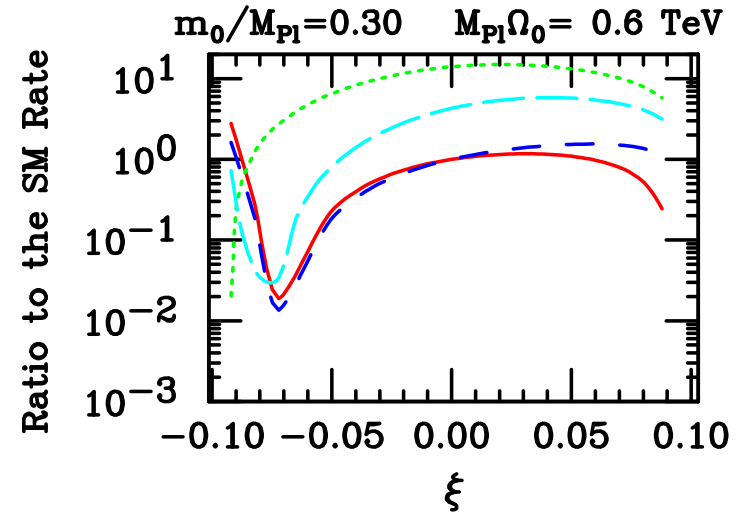
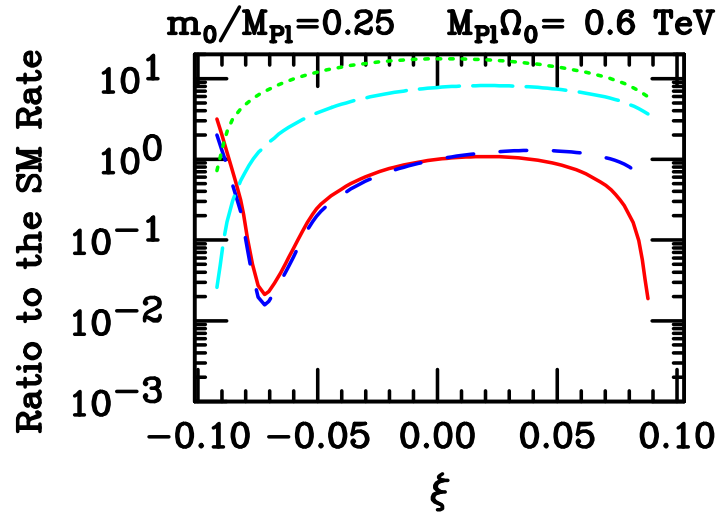
125 GeV:  $\gamma\gamma \sim 2 \times \text{SM}$  and  $4\ell \sim 1.5 \times \text{SM}$ .

137 GeV:  $\gamma\gamma \sim 2 \times \text{SM}$  and  $4\ell$  very suppressed.

- For  $\Lambda_\phi = 1$  TeV or 1.5 TeV, the reverse configuration of  $m_h = 137$  GeV and  $m_\phi = 125$  GeV is not good.



$h \rightarrow \gamma\gamma$ : solid red;  $h \rightarrow ZZ$ : blue dashes;  $\phi \rightarrow \gamma\gamma$ : green dots;  $\phi \rightarrow ZZ$ : cyan long dashes



**Figure 4:**  $\gamma\gamma$  and  $ZZ$  rates relative to SM vs  $\xi$  taking  $\Lambda_\phi$  fixed at 1.5 TeV.

## Signals at 125 GeV and 500 GeV

- Generally speaking, it is easy to find  $\Lambda_\phi = 1$  TeV or 1.5 TeV choices for  $\xi$  and  $m_0/m_{Pl}$  which give a good description of the 125 GeV signal either for  $m_h = 125$  GeV and  $m_\phi = 500$  GeV or  $m_h = 500$  GeV and  $m_\phi = 125$  GeV.
- However, it is typically the case that the  $4\ell$  signal at 500 GeV is at least as large as the SM value.

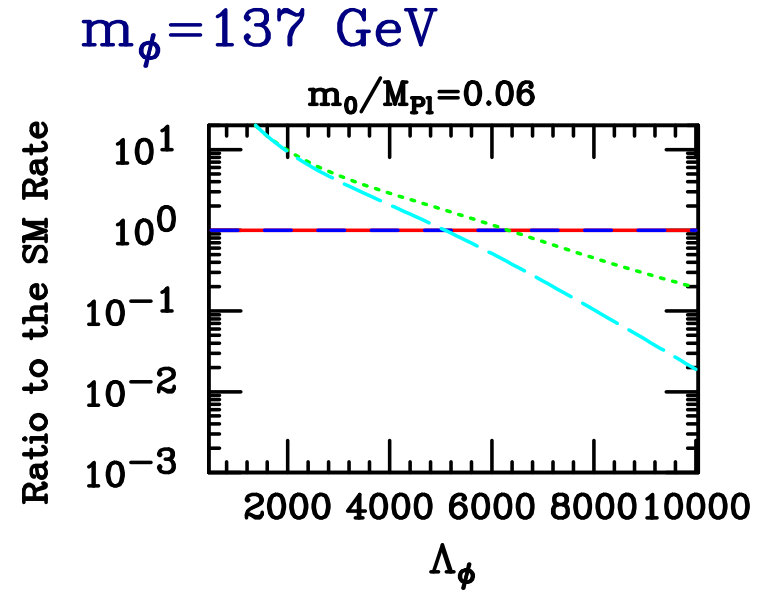
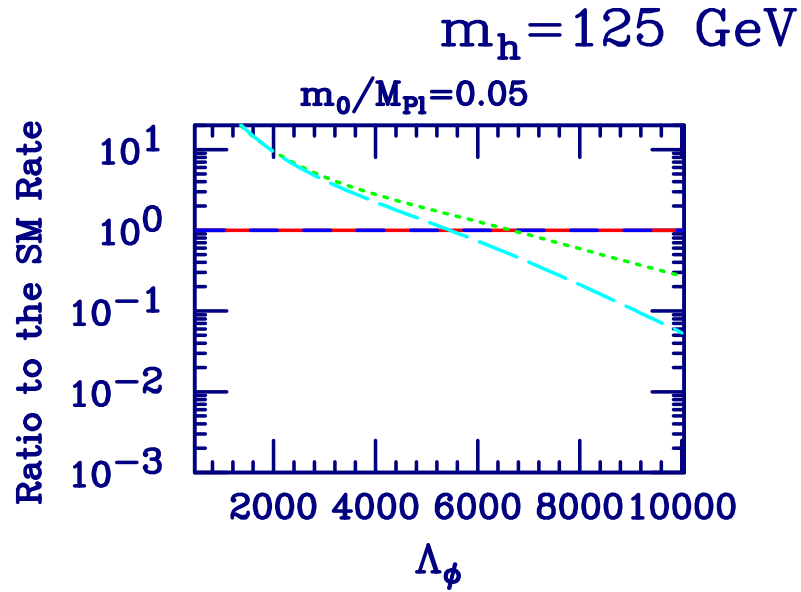
This is not in disagreement with CMS observation in this mass region, but ATLAS has a deficit.
- Thus, it may be preferred to take either the  $h$  or the  $\phi$  to have mass  $\gtrsim 1$  TeV with the other located at 125 GeV. It is then possible to describe the 125 GeV excesses while avoiding any possible conflict with the 500 GeV region data.

## SM 125 GeV signal, with varying $\Lambda_\phi$

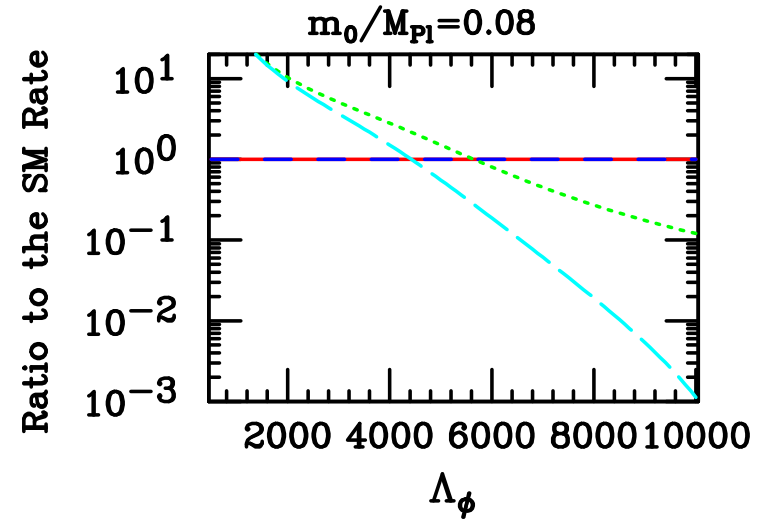
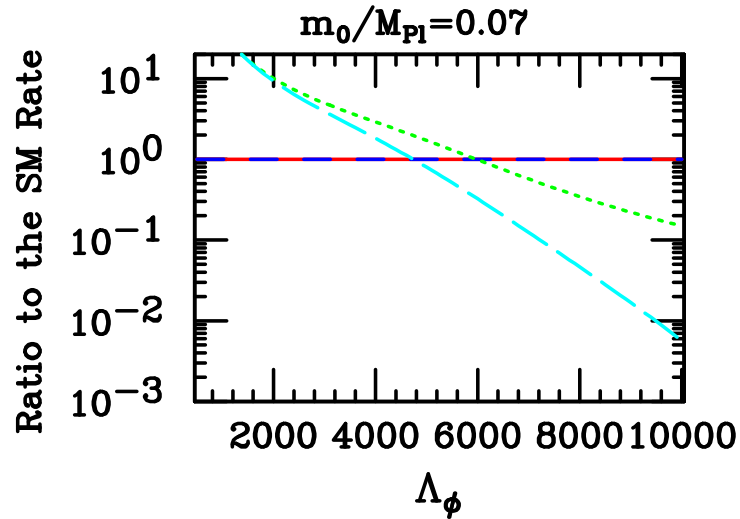
- Perhaps the signal at 125 GeV will look very precisely SM-like after more  $L$  is accumulated.

Then, one should probably take  $\xi = 0$  (no mixing) and ask what the constraints are if there is a radion at some nearby mass. We consider  $m_\phi = 137$  GeV, a signal that might survive.

- Fig. 5 shows  $\gamma\gamma > 4\ell$  at  $m_\phi$  is always the case. The unmixed radion cannot describe a  $4\ell > \gamma\gamma$  excess.
- A decent fit to the current CMS  $\gamma\gamma$  excess at 137 GeV is achieved for quite modest  $m_0/m_{Pl} = 0.05$  and  $\Lambda_\phi \sim 5.5$  TeV!



$h \rightarrow \gamma\gamma$ : solid red;  $h \rightarrow ZZ$ : blue dashes;  $\phi \rightarrow \gamma\gamma$ : green dots;  $\phi \rightarrow ZZ$ : cyan long dashes



**Figure 5:**  $\gamma\gamma$  and  $ZZ$  rates relative to SM vs  $\Lambda_\phi$  taking  $\xi = 0$ .

## Conclusions

It seems likely that the Higgs responsible for EWSB is not buried

Perhaps, other Higgs-like objects are emerging.

But, we must never assume we have un-buried all the Higgs.





Certainly, I will continue watching and waiting



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